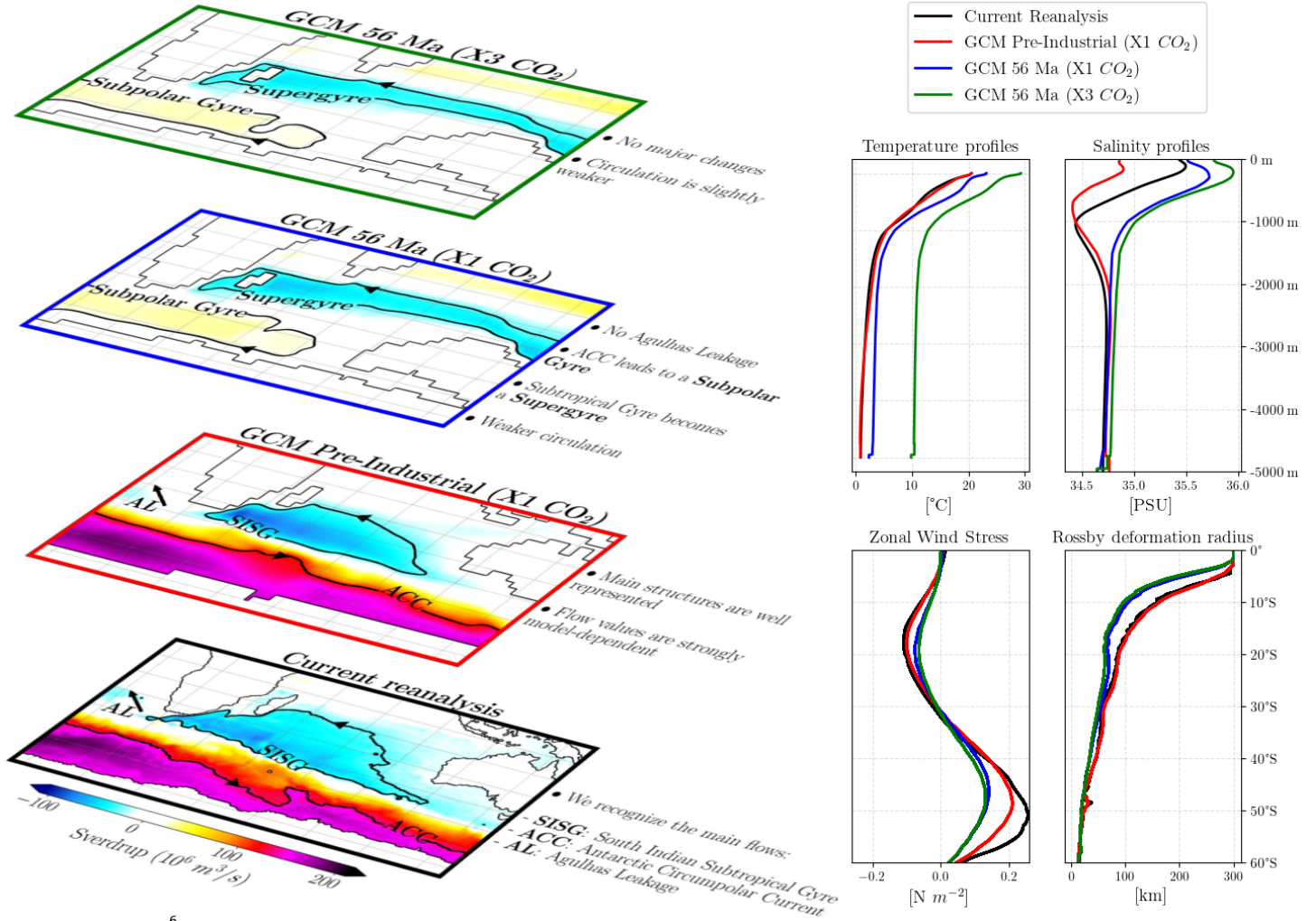


1 Graphical Abstract

2 The Early Eocene South Indian subtropical gyre in DeepMIP paleoclimate simulations

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7 Highlights

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- 12 • The Early Eocene South Indian subtropical gyre in DeepMIP paleocli-
13 mate simulations is substantially different from the present-day situa-
14 tion.
- 15 • Setting General Circulation Models in an Eocene configuration results
16 in a complete reshaping of the gyre, which weakens, extends into the
17 Pacific, and is no longer connected to the Atlantic.
- 18 • We investigate the contribution of biases associated with paleoclimate
19 models, paleogeography, and CO₂ in these changes, then compare them
20 with proxy studies.

21 The Early Eocene South Indian subtropical gyre in
22 DeepMIP paleoclimate simulations

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25 **Abstract**

The South Indian Subtropical Gyre (SISG) is a major ocean current system contributing to the global circulation. Simulations from the Deep-Time Model Intercomparison Project (DeepMIP) for the Eocene SISG reveal substantial changes in oceanic and atmospheric circulation primarily driven by shifts in paleogeography. The results indicate a reconfiguration and weakening of the gyre no longer connected to the Atlantic (collapsed Agulhas leakage) and extending into the Pacific and a southward displacement of the Subtropical Front (STF). Additionally, the simulations show significantly warmer deep waters (by 9°C), the disappearance of the freshwater nature of Antarctic Intermediate Water, and a reduced meridional temperature gradient; findings that are broadly consistent with available proxy reconstructions. This study serves as a first step in disentangling robust patterns related to the SISG from inter-model differences of varying magnitude, thereby highlighting the role of model dependency and addressing potential biases in General Circulation Models (GCM) simulating the Eocene.

26 *Keywords:*

27 South Indian subtropical gyre, Early Eocene, DeepMIP

1. Introduction

1.1. *The Indian Ocean subtropical gyre, a major element of the global oceanic circulation*

Oceanic gyres are essential for regulating global climate, heat and salinity transport, and carbon cycling. They redistribute warm and cold waters, influencing regional climates, storm patterns, and monsoons. Gyres connect wind-driven and overturning circulation, impacting ocean circulation changes and climate variability over time (Wells, 2003; Hasselmann, 1991). They also play a key role in marine ecosystems, affecting nutrient distribution to fish migration (Reid et al., 1978). The Southern Hemisphere ocean hosts a nested system of three interconnected subtropical gyres in the Atlantic, Indian, and Pacific basins, together forming a supergyre (Ridgway and Dunn, 2007; Tilburg et al., 2001). Located in the southern part of the Indian Ocean, the South Indian Subtropical Gyre (SISG) is an integral component of this larger system (Figure 1). Circulating counterclockwise, it is bounded by the South Equatorial Current (SEC) to the north (Tomczak and Godfrey, 1994), which moves along the equator toward the east coast of Africa. Once reaching the African coast, the SEC is split by Madagascar into two western boundary currents: the North Madagascar Current (NMC) flowing north and the East Madagascar Current (EMC), which flows south along the eastern coast of Madagascar (Chapman et al., 2003). After passing the northern tip of Madagascar, the NMC flows to the west and splits when reaching the African continent. Its southern branch feeds the Mozambique Channel, dominated by large eddies (Halo et al., 2014). South of the Mozambique Current, the flow merges with the EMC to form the Agulhas Current (AC), the largest western boundary current of the southern hemisphere (Lutjeharms, 2006), transporting 84 Sv (Beal et al., 2015). At the southwestern edge of the Agulhas Bank, the main part of the AC retroflects to flow eastward, giving rise to the Agulhas Return Current (ARC) (Lutjeharms and Ansorge, 2001), while its associated leakage feeds the South Atlantic Ocean. Approximately 15 Sv of warm and salty water from the AC flows into the South Atlantic through Agulhas rings (Richardson, 2007), and this may influence the Atlantic Meridional Overturning Circulation (AMOC) (Beal et al., 2011). The ARC largely recirculates within the subtropical gyre, while the remainder becomes the South Indian Ocean Current (SIOC) beyond 70°E (Lutjeharms and Ansorge, 2001). With the Antarctic Circumpolar Current (ACC), ARC and the SIOC delimit the Subtropical Front (STF) separating warm, salty Subtropical Wa-

65 ter from colder, fresher Subantarctic Water (Stramma, 1992). A part of the
 66 SIOC feeds the West Australian Current, which connects the SIOC to the
 67 SEC by flowing northward along the west coast of Australia. This defines
 68 the SISG's eastern boundary. The Indonesian Throughflow (ITF) primarily
 69 controls the connection between the Indian and Pacific Oceans. The ITF
 70 allows warm, low-salinity waters from the western Pacific to enter the Indian
 71 Ocean, influencing the dynamics and strength of the gyre (Lambert et al.,
 72 2016).

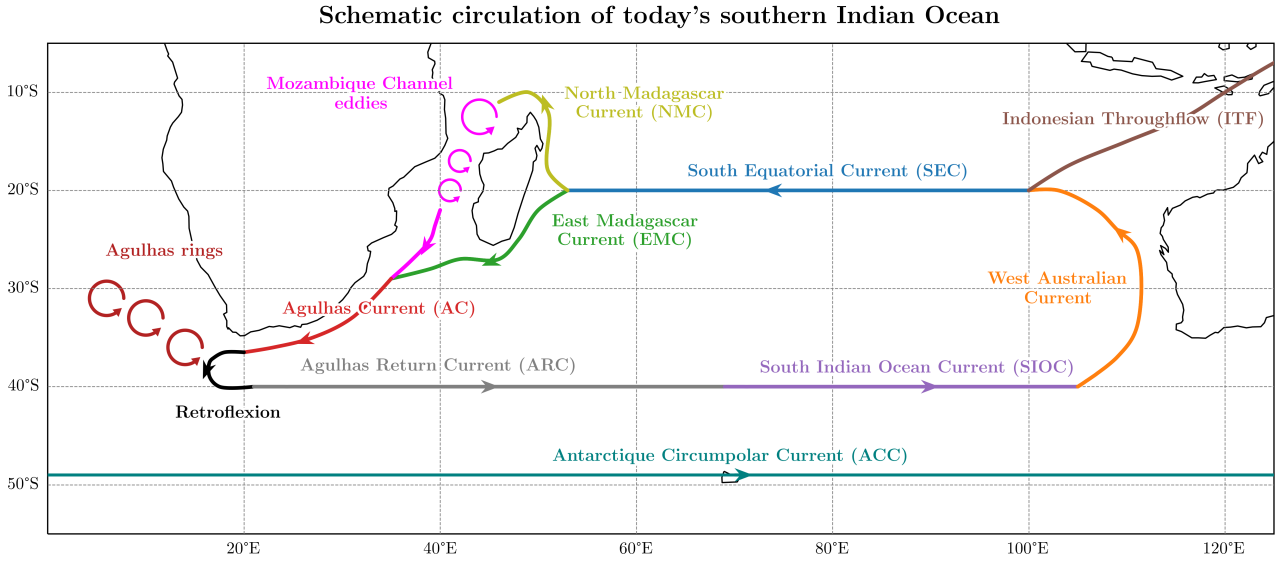


Figure 1: Schematic circulation of the southern Indian Ocean. The main currents of the SISG are represented.

73 1.2. During the Eocene, the geography, climate and ocean were different

74 At global scale, Early Eocene (56-50Ma) geography, climate and oceans
 75 were markedly different from today. With the Drake Passage and the Tas-
 76 man Gateway closed (or narrow and shallow compared to today) at the time,
 77 strong subpolar gyres were maintained in the Weddell and Ross Seas accord-
 78 ing to models and neodymium isotopes (Zhang et al., 2020; Huber et al.,
 79 2004; Cramwinckel et al., 2020). Similar to the present-day North Atlantic,
 80 these subpolar gyres transported saline waters from lower to higher lati-
 81 tudes, sustaining a strong meridional overturning circulation (~ 40 Sv at
 82 60°S) and enabling significant heat transport in the Southern Hemisphere

(Zhang et al., 2020). This mechanism contributed to the weak north-south temperature gradients and the maintenance of temperatures too high for ice sheet formation in Antarctica (Eichenseer and Jones, 2024). Surface temperatures in tropical regions during the Eocene reached 30 to 35°C, while polar regions experienced a temperate to subtropical climate with temperatures ranging from 10 to 15°C (Stokke et al., 2020; Sluijs et al., 2020). This extreme warmth is partly attributed to elevated atmospheric CO₂ concentrations, estimated to be three to six times preindustrial levels (Royer et al., 2023). The interannual proxy study by Ivany et al. (2011) using fossil bivalves and sedimentary records from the Eocene indicates quasi-periodic variations in sea surface temperature (SST), precipitation, and sea-level pressure in the equatorial Pacific, consistent with modern ENSO dynamics. Evidence in palynological assemblages suggests that the East Asian Monsoon began to emerge during the Early Eocene, driven by changes in temperature and precipitation patterns (Su et al., 2022). The absence of ice sheets resulted in sea levels up to a maximum of approximately 70 m higher than present (Miller et al., 2024), submerging extensive coastal areas and promoting the formation of epicontinental seas (Lunt et al., 2012). The hydrological cycle was intensified, creating warm and humid conditions even in mid-latitudes (Cramwinckel et al., 2023). Unlike today, where deep water formation occurs in the North Atlantic, during the Eocene, deep waters were primarily formed in the North Pacific as indicated by neodymium isotopes (McKinley et al., 2019), but confirmed only by one over 8 models in numerical studies (Zhang et al., 2022) and in the Southern Hemisphere in agreement with both models and neodymium isotopes (Zhang et al., 2022; Huck et al., 2017).

1.3. Investigating the early Eocene SISG using models in an Early Eocene configuration

These elements suggest that the SISG must have been significantly different from its present-day configuration. Recent progress in paleoclimate modeling, as demonstrated by studies using advanced Earth system models, has enabled increasingly accurate simulations of key aspects of past climate states—including global temperature distributions, precipitation regimes, and large-scale ocean circulation patterns (Lunt et al., 2020; Huber and Caballero, 2011; Zhang et al., 2022). However, these models generally operate at relatively low spatial resolution, which limits their ability to resolve mesoscale features such as oceanic eddies and narrow boundary currents. As a result, it remains uncertain whether they can accurately capture the structure and dy-

120 namics of systems like the SISG, particularly in regions where key circulation
121 features depend on mesoscale processes—for example, the eddy-dominated
122 flow through the Mozambique Channel. This limitation poses challenges for
123 assessing the realism of gyre representation in paleoclimate simulations and
124 calls for cautious interpretation of their outputs. This is the focus of our
125 investigation in this study: How was the SISG system during Early Eocene?
126 Can paleoclimate simulations provide insights? What are the robust find-
127 ings? Are they in agreement with proxy studies? Do models agree with each
128 other on these aspects?

129 To address these questions, we use multiple simulations carried out within
130 the framework of the Deep-Time Model Intercomparison Project (DeepMIP),
131 which provides a coordinated set of global coupled climate models at pre-
132 industrial and early Eocene periods. These simulations offer a valuable op-
133 portunity to investigate the structure of the SISG under markedly different
134 paleogeographic and climatic conditions. We focus on key features of the
135 SISG, analyzing its structure and intensity across the different model out-
136 puts. By comparing Eocene simulations with pre-industrial control runs and
137 current reanalysis (modern numerical model that assimilates historical ob-
138 servations), we identify robust patterns and assess inter-model differences.
139 This approach allows us to disentangle the respective contributions of model
140 biases, paleogeography, and atmospheric CO₂ concentrations. We also ex-
141 amine how these simulations align with the available proxy records from
142 the literature. Proxy reconstructions provide indirect constraints on past
143 ocean states and are subject to uncertainties related to calibration and sta-
144 tistical uncertainty in proxy-environment relationships (McClelland et al.,
145 2021). Moreover, proxy records are spatially discrete and may be representa-
146 tive only of local or regional processes, which complicates their extrapolation
147 to basin-scale circulation features resolved by climate models. Additionally,
148 preservation and post-depositional alteration (diagenesis) of proxy carriers
149 can modify original geochemical signals and introduce further uncertainty
150 (Edgar et al., 2015). These uncertainties, combined with potential issues
151 of spatial representativeness and age control in sedimentary archives, mean
152 that many proxy records integrate climate signals over time windows that
153 do not match the model averaging strategy. In all models, the SISG ex-
154 hibits a markedly different configuration compared to today: the gyre is
155 weaker, extends into the Pacific, with different water masses composition
156 and a collapsed Agulhas leakage. These results highlight the dominant role
157 of paleogeography in shaping oceanic circulation during the Early Eocene.

158 They also provide a foundation for identifying systematic biases in paleocli-
159 mate models and for refining future paleoceanographic reconstructions of the
160 Indian Ocean.

161 2. Models

162 DeepMIP is a working group within the PMIP4 Paleoclimate Modelling
163 Intercomparison Project, itself part of the sixth phase of the Coupled Model
164 Intercomparison Project, CMIP6 (Eyring et al., 2016). DeepMIP aims to
165 enhance our understanding of "Deep Time" climates through international
166 collaboration. This initiative focuses on geological periods with atmospheric
167 CO₂ concentrations comparable to those projected for the late 22nd cen-
168 tury—levels that have no recent equivalent in modern climate archives. In-
169 deed, while current climate models are tested against modern-era observa-
170 tions, they must also be validated using past periods characterized by sim-
171 ilarly high CO₂ concentrations. DeepMIP is an excellent opportunity to
172 compare the SISG in various Early Eocene simulations. However, paleocli-
173 mate modelling is not without bias. One of the primary limitations is the
174 resolution of the models. The thousands of years of simulations required to
175 reach equilibrium in coupled GCMs for paleoclimate simulations necessitate
176 the use of low-resolution grids ($> 1^\circ$ for ocean models) due to computational
177 cost constraints. Nevertheless, fine-scale structures (10-100 km) play a cru-
178 cial role in ocean circulation, particularly in terms of transport processes,
179 fronts, and mixing layers (Hirschi et al., 2003; Li et al., 2024). This is espe-
180 cially true for the representation of western boundary currents, such as the
181 AC (Schwarzkopf et al., 2019; Holton et al., 2017). For instance, the Agulhas
182 leakage, which connects the Indian Ocean to the Atlantic, occurs through
183 Agulhas rings: eddies that require at least mesoscale resolution to be repre-
184 sented in models (Schubert et al., 2019). The initial conditions and forcings
185 used in simulations (such as greenhouse gas concentration, paleotopography,
186 or ice distribution) often rely on indirect reconstructions, which may contain
187 uncertainties or errors. Convection and vertical mixing are not parameter-
188 ized in the same way across different models. Paleoclimatic reconstructions
189 (isotopic ratios, foraminifera, etc) involve uncertainties or may lack data de-
190 pending on the time periods or regions. The numerical schemes used in
191 models can overly smooth gradients (e.g., temperature, salinity), and so re-
192 ducing the accuracy of simulations. Interactions between model components
193 (ocean, atmosphere, ice, biogeochemistry) can introduce errors when the cou-

plings are not properly optimized. The nine climate models in the DeepMIP and the Eocene simulations are documented in detail by Lunt et al. (2020). The Eocene paleogeography and bathymetry used in all DeepMIP simulations follow the global reconstructions of Herold et al. (2014), which provide a standardized representation of continental positions, ocean gateways, and seafloor depth for the early Eocene (55 Ma). These reconstructions constitute the common boundary conditions adopted by DeepMIP models (except for NorESM). Zhang et al. (2022) supplement provides additional insights about diapycnal mixing parametrization, bathymetry boundary condition and initial thermohaline conditions. In addition to the various CO₂ experiments, all modeling groups were required to carry out a pre-industrial (PI) simulation for comparison purposes, which was to be as close as possible to the CMIP6 standard PI-Control simulation (Eyring et al., 2016).

Model	Laboratory	CO ₂	(lat × lon), z levels	Ocean Grid	Run length (yr)
CESM1.2-CAM5	NCAR (USA)	x1, x3, x6, x9	(1° × 1°), 60	Greenland grid (polar shifted)	2000–2600
COSMOS-landveg-r2413	MPI-M (Germany)	x1, x3, x4	(3° × 1.8°), 40	Bipolar curvilinear	9500
GFDL-CM2.1	NOAA-GFDL (USA)	x1, x2, x3, x4, x6	(1° × 1.5°), 50	Tripolar	6000
HadCM3B-M2.1aN	Met Office Hadley Centre (UK)	x1, x2, x3	(1.25° × 1.25°), 19	Cartesian	7400
HadCM3BL-M2.1aN	Met Office Hadley Centre (UK)	x1, x2, x3	(3° × 2.5°), 20	Cartesian	7500
INMCM-CM4-8	INM (Russia)	x6	(1° × 0.5°), 40	Generalized spherical (C-grid)	1150
IPSLCM5A2	IPSL (France)	x1.5, x3	(2° × 2°), 31	Tripolar	4000
MIROC4m	CCSR/NIES/FRCGC (Japan)	x1, x2, x3	(1° × 1°), 44	Regular grid with tropics refinement	5000
NorESM1-F	Norwegian Climate Center (Norway)	x2, x4	(1° × 1°), 46	Tripolar	2100

Table 1: Summary of paleoclimate models participating in DeepMIP: laboratory affiliation, experiments performed, ocean grid resolution and type, and simulation length.

Eocene boundary conditions are significantly different from those of the present day, implying that these simulations initialized with crude ocean thermohaline structures will require several millennia to reach an equilibrium state (Kennedy-Asser et al., 2020). Over the last 1000 years of simulation, deep ocean temperature trends are less than 0.5°C.

The first step of this study will be to determine robust patterns in the low-resolution modeling of GCMs by comparing DeepMIP PI outputs to state-of-

the-art reanalyses and climatologies, such as GLORYS (Lellouche et al., 2018) for ocean circulation, the World Ocean Atlas 2023 (WOA) for water mass properties at 1/4° resolution (Reagan et al., 2024) and the European Centre for Medium-Range Weather Forecasts (ECMWF) Reanalysis v5 (ERA5) for atmospheric conditions (Hersbach et al., 2020).

3. Methods

In this study, we analyzed the SISG of seven DeepMIP GCMs. These models include CESM1.2 (hereafter CESM), COSMOS-landveg-r2413 (hereafter COSMOS), GFDL-CM2.1 (hereafter GFDL), IPSL-CM5A2 (hereafter IPSL), MIROC4m (hereafter MIROC), HadCM3B and HadCM3BL. For each of these models, we will compare three configurations: the pre-industrial configuration (hereafter PI) with a CO₂ level of 280 ppm, an early Eocene configuration with the same CO₂ level as PI (hereafter paleo x1 CO₂), and an early Eocene configuration with three times the CO₂ level of paleo x1 CO₂, i.e., 840 ppm (hereafter paleo x3 CO₂). We do not include INMCM or NorESM, as neither has a paleo x3 CO₂ configuration. Additionally, for wind stress analysis, we exclude COSMOS as it does not provide wind stress as an output. The outputs provided by DeepMIP consist of monthly climatologies (averaged over the last 100 years of each simulation) and monthly time series for the last 100 years. The variables are described and made available as open access (Steinig et al., 2024). Given the differences in grid configurations, we pre-processed the model outputs to a higher resolution to preserve maximum information. Temperature and salinity vertical sections are re-interpolated linearly on grids with a 0.5° latitudinal resolution and 5 m vertical resolution. We analyzed the SISG within a domain covering the southern part of the Indian Ocean (70°S–0°S, 0°E–150°E) for the following variables: wind stress curl, barotropic streamfunction, STF, water masses, and Rossby deformation radius, as detailed in the next subsections.

3.1. Mean Wind Stress Curl

Wind Stress Curl is the forcing factor of the Sverdrup balance (Sverdrup, 1947):

$$\beta \frac{\partial \psi}{\partial x} = \frac{\nabla \times \boldsymbol{\tau}}{\rho}$$

where $\beta = \frac{\partial f}{\partial y}$ (s⁻¹ m⁻¹) is the meridional gradient of the Coriolis parameter f (s⁻¹), ψ (m³ s⁻¹), hereafter expressed in Sverdrups with 1 Sv = 10⁶ m³ s⁻¹

247 is the barotropic streamfunction (with $V = \frac{\partial \psi}{\partial x}$ the vertically integrated
 248 meridional transport, in $\text{m}^2 \text{s}^{-1}$), ρ is the seawater density (kg m^{-3}), and
 249 $\nabla \times \boldsymbol{\tau} = \frac{\partial \tau_y}{\partial x} - \frac{\partial \tau_x}{\partial y}$ is the wind stress curl (N m^{-3}), the direct driver of
 250 oceanic gyres. Position and intensity of oceanic gyres can vary on geological
 251 timescales. Zhang et al. (2025) demonstrated that this is the case for the
 252 North Pacific Gyre, which shifted northward at the onset of the Eocene, pri-
 253 marily driven by wind-driven transport. In our case, the computation of the
 254 curl requires evaluating τ_x and τ_y on the sides of each grid cell. This implies
 255 that land masks are slightly modified, which can impact the calculation more
 256 significantly at lower resolutions (for example, it leads to a closed Mozam-
 257 bique Channel for HadCM3BL). The zonally-averaged wind stress curl is
 258 calculated over the region spanning 50°E to 100°E , away from continental
 259 masses.

260 *3.2. Mean barotropic streamfunction, Agulhas leakage, gyre recirculation and* 261 *ACC transports*

262 The barotropic streamfunction ψ represents the depth-integrated mean
 263 horizontal flow and is a key diagnostic for identifying the structure and
 264 strength of large-scale ocean gyres. It is derived from the curl of the barotropic
 265 velocity field, ensuring mass conservation and representing the circulation
 266 around islands and continental boundaries. In this study, the streamfunc-
 267 tion is computed by following a Helmholtz decomposition by solving a Poisson
 268 equation using a method that incorporates both a particular solution for the
 269 domain and homogeneous solutions for individual islands. The final solution
 270 is constructed to ensure that the circulation around each island matches that
 271 imposed by the barotropic transport field. This method, based on the frame-
 272 work of Colin de Verdière and Ollitrault (2016) and solved via successive
 273 over-relaxation (Yang and Matthias, 2007), is detailed in the first section
 274 of the appendix document. Because the DeepMIP simulations provide low-
 275 frequency (climatological) output, the Agulhas leakage is diagnosed here by
 276 identifying streamlines of ψ that connect the Indian and Atlantic Oceans un-
 277 der specific geometric and directional criteria rather than using Lagrangian
 278 methods (Richardson, 2007). These include crossing key transects (e.g., the
 279 Good Hope Line and 40°E) and flowing from east to west north of the STF.
 280 This method is detailed in the second section of the appendix. By applying
 281 this method to the annually-averaged GLORYS horizontal transport, we ob-
 282 tain 13.1 Sv, which is close to the value derived from the Lagrangian method
 283 applied to observations (Richardson, 2007) and from eddy-resolving simula-

284 tions (Schmidt et al., 2021). The subtropical gyre circulation is defined as
 285 the difference between the maximum of ψ in the center of the recirculating
 286 subgyre and the Agulhas leakage. This gives 69.8 Sv for GLORYS which is
 287 close to the 66 Sv from Dijkstra and De Ruijter (2001). For the ACC, we
 288 take the transport value closest to Antarctica (this value should be at first
 289 order independent of longitude). This gives a value of 151.2 Sv for GLORYS
 290 which is slightly lower than the recent measure at the Drake passage 173.3
 291 ± 10.7 Sv (Donohue et al., 2016).

292 *3.3. STF position and water masses*

293 To study the STF position, which marks the southern boundary of the
 294 gyre and separates warm subtropical waters from cold subpolar waters, we
 295 define temperature and salinity sections across the Mozambique Channel.
 296 Additionally, the STF plays a crucial role in the connection between the
 297 SISG and the Atlantic Ocean (Dencausse et al., 2011) which makes it in-
 298 teresting for us to study. Salinity sections also provide insight in the repre-
 299 sentation of water masses. For the temperature sections, we computed the
 300 SST difference along the section between the two latitudes 60°S and 15°S at
 301 10 m depth (noted $\Delta_y \text{SST}$). In paleoclimate simulations, the characteristic
 302 salinity and temperature values of water masses can differ significantly from
 303 those of the present day. Therefore, the method for identifying the STF de-
 304 scribed by Belkin et al. (1998) cannot be applied. Instead, the position of
 305 the STF is determined as the latitude corresponding to the maximum merid-
 306 ional temperature gradient along the section at 10 m depth. Vertical profiles
 307 of temperature, salinity, and density averaged over the SISG region provide
 308 insights into water mass characteristics. We analyze water mass composition
 309 using Θ - S diagrams computed over 20°S–40°S and 50°E–100°E, interpolated
 310 onto a common z-grid to allow envelope representation. Stratification de-
 311 fined by the square of the buoyancy frequency N^2 (s^{-2}), is examined over
 312 the upper 750 m, where most variations occur. This approach helps trace
 313 the origin of water masses within the gyre and assess how they respond to
 314 different experimental conditions.

315 *3.4. Rossby deformation radius (R_0)*

316 Finally, we examine the typical size of the oceanic mesoscale eddies that
 317 would develop through the instability of large scale currents if the model
 318 resolution were sufficient. As conveyors of heat, salt, momentum and marine
 319 properties, mesoscale eddies are dominating the present oceanic circulation

(Chelton et al., 2007; Ferrari and Wunsch, 2009). As the mesoscale eddy length scales have been observed to align with the Rossby deformation radius (R_0 hereafter expressed in km) (Stammer, 1997), this parameter should give first insights of the mesoscale eddy properties of the past (even in non eddy resolving simulations). This length scale depends on the ocean depth and stratification, and the local rotation of the Earth (the Coriolis parameter). For this purpose, we compute the Rossby deformation radius, which defines the length scale of baroclinic instability, directly influencing eddy size and baroclinic wave propagation (Chelton et al., 1998). Rossby deformation radius is directly affected by the large changes in bathymetry and stratification between paleo and pre-industrial simulations. Rossby deformation radius is obtained by solving the eigenvalue problem for the vertical structure of perturbations horizontal velocities (λ_n^2 , with the associated Rossby deformation radius $R_n = \lambda_n^{-1}$ (Chelton et al., 1998)):

$$\partial_z \left(\frac{f^2}{N^2} \partial_z F \right) + \lambda_n^2 F = 0$$

with surface and bottom boundary conditions $\partial_z F = 0$ at $z = 0$ and $z = -H$. In this formula, the Brunt-Vaisala frequency N (defined as $N^2 = -\frac{g}{\rho_0} \partial_z \rho$) is computed using the potential density ρ (kg m^{-3}) calculated with an empirical relationship (McDougall et al., 2009) as a function of temperature and salinity, which are interpolated level by level onto a common 3D z-grid (1000 regularly spaced vertical levels).

4. Results

4.1. Wind Stress Curl: A shifted and weakened wind circulation system

By controlling Sverdrup transport, wind stress curl directly drives the large-scale gyres. It plays a key role in their development and maintenance by inducing horizontal pressure gradients that sustain geostrophic flows (Sverdrup, 1947). Positive wind stress curl (Figure 2a, red orange and yellow areas) are associated with equatorward meridional velocities in the interior ocean and poleward return current along the western boundary, driving an anticyclonic circulation in the southern hemisphere, typical of subtropical gyres. At present day, trade winds (with a maximum around 17°S in ERA5) are driven by the Hadley cell which drives tropical circulation with rising air at the equator and sinking at the subtropics, while westerly winds (peaking

around 51°S in ERA5) are driven by the Ferrel cell, the mid-latitude over-
turning air circulation that moves in the opposite direction of the Hadley cell
(Fairbridge, 2005). Their maxima are identified using the zonally-averaged
zero wind stress curl. With varying wind stress intensity and westerly wind
positions (Figures 2b, c), the PI outputs produce a tripolar wind circulation
system. Theoretical frameworks link the poleward boundary of the subtropi-
cal gyres to the latitude where the wind stress curl is zero (De Ruijter, 1982).
The position and intensity of oceanic gyres vary on geological timescales.

The most robust results across the model runs from this figure are the
weakening of the entire wind circulation system and the northward shift of
the westerlies in the Eocene simulations (Figure 2a, columns X1 and X3).
Models unanimously show a decrease in the zonal wind stress, with an av-
erage reduction of $\sim 50\%$ (Figure 2c). This is consistent with Zhang et al.
(2022), who found a lower wind stress in the x3 configuration compared to
the PI configuration for the IPSL model in the Southern Hemisphere. This
is also the case for other models, even in the paleo x1 CO₂ configurations.
This indicates that CO₂ is not the only driving force behind the weakening
of wind stress. This phenomenon is explained by the past continental config-
uration, which enhances meridional heat transport (Zhang et al., 2022) and
tends to reduce the meridional temperature gradient (a weaker meridional
temperature gradient is observed and discussed later), which plays a key role
in atmospheric circulation (Chelton et al., 2001). A decrease in wind stress
curl is also observed in the Pacific part of the supergyre (Zhang et al., 2025),
and this change is associated with Australia, which forms an orographic ob-
stacle in the Eocene simulations. The Eocene x3 CO₂ simulations show a
northward shift of the westerlies in all models (3.6° on average compared to
ERA5). This shift is even more important in the paleo x1 CO₂ simulation
(5.2°) and is attenuated by the increased CO₂. A less robust result is the
southward shift of the trade winds (1.4°). However, this shift is more ro-
bust (and important) in the paleo x1 CO₂ simulations (2.3°). As a result,
we also note that the Eocene simulations show better intermodel agreement
compared to the PI simulations.

4.2. Barotropic streamfunction: A reshaped and weaker gyre system

The barotropic streamfunction ψ is primarily driven by the wind stress
curl at the center of the gyre but is strongly influenced by other dynamical
processes such as bottom pressure torque or inertial and viscous effects in
the case of a western boundary current like the AC (Penven et al., 2021).

389 As ψ is defined relative to an arbitrary additive constant (Figure 3a), we
 390 set $\psi = 0$ on the African continent. Positive/negative values mean anti-
 391 clockwise/clockwise integrated transport relative to Africa. For example,
 392 AC (clockwise compared to Africa) leads to negative values. Figure 3b sum-
 393 marizes the value of the Agulhas Leakage, the ACC, the subpolar gyre and
 394 gyre recirculation. When applied to GLORYS reanalysis, the method gives
 395 values close to those found in the literature. When applied to the PI sim-
 396 ulations, it gives strongly varying results: Agulhas Leakage is ranging from
 397 5.2 Sv for IPSL to 63.6 Sv for CESM. This may be attributed to several
 398 factors, such as the position of the STF. ACC transport values range from
 399 101.5 Sv for IPSL to 172.9 Sv for GFDL, we did not prospect the origin of
 400 these variabilities, whose transport is controlled by numerous factors such as
 401 the westerly wind stress, density gradients, eddy activity, and bathymetry
 402 (Lamy et al., 2024). The inter-model spread in subgyre recirculation is also
 403 high, ranging from 16.6 Sv for CESM to 76.0 Sv for MIROC. This should be
 404 linked to the overestimation of the Agulhas leakage because if we look at the
 405 maximum of the SISG (the AC transport), only IPSL and HadCM3BL are
 406 significantly lower than GLORYS.

407 Compared to the PI configurations, the Early Eocene simulations show
 408 radical changes in the SISG and the Indian Ocean circulation in general.
 409 Across all models, SISG is no longer exclusively connected to the South Pa-
 410 cific via the ITF and Tasman Leakage. Instead, part of the flow expands
 411 across the entire Pacific (Fig. 3a), passing north of Australia and merging
 412 with the South Pacific Subtropical Gyre to form a "supergyre", following
 413 the formulation used by Zhang et al. (2025) to describe this connection. A
 414 portion of the flow continues to recirculate within the Indian Ocean. This
 415 recirculating component can be quantified by subtracting the Pacific outflow
 416 from the total retroflexion. Here, we do not explicitly represent the Pacific
 417 portion of the gyre. This component is illustrated by Zhang et al. (2025).
 418 In all simulations, mid-latitudes transport between the Indian and Pacific
 419 Oceans is enhanced under this new configuration. A subtropical cyclonic
 420 gyre takes place east of India in all Eocene simulations, the ACC no longer
 421 exists, and a cyclonic subpolar gyre is developing in the Austral region with
 422 a flow ranging from 31.5 Sv for HadCM3BL to 53.0 Sv for IPSL. This gyre
 423 is carrying water from mid to high latitudes and thus, contributes to keep a
 424 weak meridional temperature gradient. This should maintain a much weaker
 425 SISG (especially considering the absence of leakage). The AC (Agulhas leak-
 426 age + Gyre recirculation) is ranging from 15.9 Sv for HadCM3BL to 56.9 Sv

427 for MIROC in the x3 simulation (CO_2 also contributes to this weaker circu-
428 lation) against an average of 74.4 Sv for the PI simulations and 73.2 Sv for
429 GLORYS. In all Eocene simulations, the boundary between the subtropical
430 and the subpolar gyre (shown by the $\psi = 0$ contour in Fig. 3a) connects to
431 the southern tip of Africa, indicating the absence of Agulhas leakage. This is
432 largely due to the southward shift of the position of Africa now intersecting
433 the STF. Indeed, Africa is much further south ($\sim 10^\circ$) than the average STF
434 poleward shift (3.7° in average in the paleo x1 CO_2 simulations).

435 4.3. *Water masses temperature and salinity in the DeepMIP simulations*

436 To understand the thermohaline changes in the SISG region, we analyze
437 the water masses represented in the different DeepMIP simulations. These
438 are traditionally explored using Θ -S diagrams, where potential temperature
439 Θ replaces depth on the vertical axis and salinity S is on horizontal axis.
440 Figure 4 shows these diagrams for the central subtropical gyre (20°S – 40°S
441 and 50°E – 100°E) for each model and configuration (PI, paleo x1 CO_2 , and
442 paleo x3 CO_2). In addition to Θ -S diagrams, vertical profiles of temperature,
443 salinity, and density are shown in Figures 5c, 6b and 6c for the same re-
444 gion. These reveal the vertical structure and variability across models. The
445 meridional SST gradients are summarized in the radial scatter plot in Fig-
446 ure 5b. Figures 5a and 6a present meridional sections of temperature and
447 salinity along the Mozambique Channel (42°E for the present configurations
448 and 37°E for the paleo configurations).

449 We identify in figure 6 the different water masses of the western part
450 of the Indian Ocean in the WOA salinity section, which are not the same
451 as the ones identified in the middle of the gyre. The Subtropical Surface
452 Water (STSW) occupies the uppermost layer of the water column, typically
453 found above 200 m depth. It is characterized by relatively high salinity due to
454 evaporation higher than precipitation in subtropical regions. It is bordered at
455 south by the STF. Red Sea Water (RSW) is present at intermediate depths,
456 usually between 900 and 1200 m. This water mass originates from the Red
457 Sea and is known for its high salinity and temperature. AAIW is found at
458 depths around 1000 m. It is characterized by lower salinity compared to
459 surrounding water masses, as it originates from the colder, fresher waters
460 around Antarctica. North Atlantic Deep Water (NADW) is found at greater
461 depths, typically between 2200 and 3500 m. This water mass is formed in
462 the North Atlantic and is characterized by relatively high salinity and oxygen
463 content. North Indian Deep Water (NIDW) is generally found in the deep

464 layers of the Indian Ocean and may be present in the northern parts of
 465 the Mozambique channel (Sætre and Da Silva, 1984). Harms et al. (2019)
 466 identified water masses in the center of the gyre using CTD measurements
 467 in our study region and found a structure similar to that given by WOA
 468 (Figure 4). In order of increasing temperature, these include Circumpolar
 469 Deep Water (CDW), present at depth and mainly in the southern part of the
 470 region, NIDW, AAIW, Subantarctic Mode Water (SAMW) located between
 471 isopycnals 1034.5 and 1035.5 kg m^{-3} and STSW.

472 DeepMIP PI models fail to accurately represent the water masses within
 473 the exact salinity range of their counterparts (differences with WOA) in the
 474 western part of the basin, they still manage to capture the structures shapes
 475 (Figure 6a). Specifically, the AAIW subducts beneath the STSW. The in-
 476 termediate waters originating from the northern part of the channel (RSW)
 477 is effectively saltier than the AAIW coming from the south. This result is
 478 also observed when comparing the models in the center of the gyre. The Θ -S
 479 diagrams successfully capture the general shape of the water masses (convex,
 480 then concave, then convex when moving from low to high density), except
 481 for HadCM3B (convex then concave). IPSL, in particular, closely matches
 482 WOA. GFDL and MIROC are also sufficiently close to WOA to represent
 483 deep water. Down to 1000 m depth (over 5°C, Figure 5d), the salinity and
 484 temperature profiles exhibit a mean bias that is respectively negative and
 485 positive compared to WOA. This results in a negative bias in the density
 486 profile (on average across all models). This pattern is also observed in the T-
 487 S diagrams, where below the isopycnal 1035 kg m^{-3} all models except IPSL
 488 are negatively shifted in salinity relative to WOA. Salinity profiles exhibit
 489 significant inter-model variability at the surface and subsurface, depending
 490 on whether they advect the STSW southward or not. There is a significant
 491 inter-model spread in terms of isotherm depth (Figure 5a). Examining the
 492 4°C isotherm in the MC section it reveals differences between the models:
 493 GFDL and COSMOS 4°C isotherm is way deeper (~ 2500 m) than WOA
 494 (~ 1500 m). The North-South temperature gradients at a depth of 10 m are
 495 slightly weaker in the DeepMIP simulations compared to WOA, with a dif-
 496 ference ranging from 0.1°C lower for HadCM3BL to 4.7°C for GFDL (Figure
 497 5b). This could contribute to explain the weaker westerlies system observed
 498 on figure 2. The temperature/salinity averaged across the models align well
 499 with WOA, with an inter-model spread of less than 1°C/0.2 PSU, which is
 500 reflected in the density profiles.

501 4.4. Warmer and saltier subtropical gyre with no AAIW during Eocene

502 Overall, the paleo configurations lead to a warmer and saltier SISG.
503 On average, across all models, the temperature shift is approximately 3°C
504 throughout the water column for paleo x1 CO₂ and around 9°C for x3 (Fig-
505 ure 5c). Regarding salinity, the shift is primarily driven by changes in geog-
506 raphy and occurs mainly in shallow waters (<1500 m in the profiles). This
507 could be explained by the SISG being much more extended, stretching east-
508 ward and reaching north of Australia up to South America (Zhang et al.,
509 2025). As a result, the SEC propagates over thousands of kilometers at 20°S,
510 allowing it ample time to warm up and accumulate high salinity through
511 evaporation before forming the STSW (as observed in the salinity sections).
512 Deep waters properties are affected by a reorganization of the location of
513 deep water formation towards the Southern Ocean. The closed Drake and
514 Tasman gateways promoted a *salt-advection feedback mechanism*, similar to
515 the present-day overturning circulation of the North Atlantic, which sus-
516 tained the poleward advection of saline subtropical water to the convective
517 regions of the Southern Ocean (Zhang et al., 2020). This salinity shift at
518 the sea surface and subsurface compensates for the warmer temperature in
519 terms of seawater density down to 2000 m. The influence of these changes
520 on density profiles highlights the α -behavior (temperature influence density
521 more than salinity) of the ocean in these simulations, which is consistent
522 with present-day conditions (Stewart and Haine, 2016). The Θ -S diagrams
523 (Figure 4) and salinity sections (Figure 6a) show that the freshwater peak
524 associated with AAIW is only represented by GFDL in the paleo x1 CO₂ and
525 paleo x3 CO₂ configurations. However, while the surface waters are saltier
526 in paleo x1 CO₂ and paleo x3 CO₂ configurations, deep water salinity is not
527 significantly affected by the paleo configuration (except for CESM, Figure
528 6a) in contrast with deep water temperature, which reaches extremely high
529 values up to 12.5°C at 4500 m (Figure 5c). This results in a linear rela-
530 tionship between temperature and salinity at a density greater than 1034
531 kg m⁻³ and a change in the Θ -S curve shape from convex-concave-convex
532 to convex-linear (it is less clear for CESM and remaining convex-concave
533 for GFDL). A final key result, previously mentioned and carrying impor-
534 tant implications for wind circulation and consequently for precipitation, is
535 the meridional temperature gradient. This gradient, associated with changes
536 in oceanic transport and the transformation of the ACC into a polar gyre,
537 decreases markedly from 11°C in CESM to 4.1°C in COSMOS.

538 4.5. Poleward shift of the STF

539 Tightly associated with the southern boundary of the subtropical gyre,
540 the location of the STF provides a synthetic indicator of the ocean dynam-
541 ics. We first compare temperature sections (as a function of latitude and
542 depth) across the Mozambique Channel from 60°S to 5°S for each of the
543 seven DeepMIP models in the three configurations (PI, paleo x1 CO₂, and
544 paleo x3 CO₂), and for the WOA (1/4°) realistic reference (Figure 5a). The
545 STF positions relative to those from WOA are calculated according to the
546 method described in section 3 and summarized in the radial scatter plot in
547 Figure 5b. DeepMIP PI simulations place the STF between 37.5°S and 43°S,
548 which is consistent with the 40.5°S position from WOA salinity and temper-
549 ature sections (Figures 5a and 6a). Except for IPSL, the other PI-models
550 underestimate surface salinity values north of the front, making it difficult
551 to visualize in these sections (Figure 6a). In the configurations paleo x1
552 CO₂, on the temperature sections, the STF shifts southward, ranging from a
553 1.5° displacement for HadCM3B to 6.5° for MIROC. Except for CESM and
554 HadCM3B, the front’s position also appears in the salinity sections, with a
555 significantly more pronounced meridional SSS gradient than in the PI con-
556 figuration, particularly for IPSL, HadCM3BL, and COSMOS. CO₂ impact
557 on the front position is not robust across all models.

558 4.6. Rossby deformation radius

559 Figure 7a represents the Rossby deformation radius to investigate how
560 the GCMs, Eocene geography and CO₂ concentrations may influence eddy
561 size, wave propagation and length scale of baroclinic instability. Indeed, near
562 the Agulhas Bank, the current’s momentum overcomes topographic vorticity
563 constraints when its width matches the Rossby deformation radius, causing
564 it to detach and retroflect (Lutjeharms et al., 2003). Rossby deformation
565 radius governs the kinetic energy cascade from the mean current to eddies.
566 Eddies with diameters close to ~40 km dominate the Agulhas Return Cur-
567 rent, contributing to its intense mesoscale variability (Tedesco et al., 2022).
568 Hutchinson et al. (2018) shows that seasonal phasing of the Agulhas trans-
569 port is found to be highly sensitive to baroclinic Rossby waves propagation
570 speeds, depending directly on Rossby deformation radius. Stratification pro-
571 files (Fig. 7c) and the zonally-averaged Rossby deformation radius in the
572 central part of the basin show that, on average, the PI simulations accu-
573 rately represent Rossby deformation radius consistent with GLORYS. The

574 differences observed are mainly due to grid differences between models which
575 once interpolated, result in varying bathymetries.

576 In paleo-simulations, significant changes in terms of Rossby deformation
577 radius should be considered relative to the position of the continents, which
578 have shifted significantly (along with the value of the Coriolis parameter)
579 between the PI configurations and the paleo simulations. For example, if we
580 consider the northern region of Madagascar, where large anticyclonic eddies
581 form (Halo et al., 2014) and currently have a strong influence on the Agulhas
582 system, values strongly decrease from ~ 150 km in PI simulations to ~ 75
583 km in paleo simulations (Figure 7a). Similarly, in the retroflexion area at
584 the southern tip of South Africa, which is particularly sensitive to mesoscale
585 activity, Rossby deformation radius strongly decreases from ~ 50 km in PI
586 simulations to ~ 20 km in paleo simulations. If we examine changes at a con-
587 stant latitude in the SISG (Figure 7b), particularly at low latitudes, Rossby
588 deformation radius is lower in paleo configurations (at 10°S , PI simulations
589 yield an average $R_0 \sim 165$ km compared to $R_0 \sim 100$ km in paleo x1 CO_2
590 and paleo x3 CO_2). This is mainly due to changes in the bathymetry used,
591 which is shallower in the Eocene configuration (Herold et al., 2014). Indeed,
592 if we analyze the stratification profiles, the SISG is generally more stratified
593 in paleo simulations, particularly in surface waters (Figure 7c).

594 5. Discussion

595 5.1. Representation of the modern SISG with the DeepMIP models

596 The shape, intensity, and water mass composition of the South Indian
597 subtropical Gyre in current reanalysis differ significantly from those in the
598 Eocene DeepMIP simulations. Main results are summarized in Figure 8:
599 changes in the shape of the SISG (panels a, b, c, and d), changes in the
600 intensity of gyre circulation (panels e and f), changes in the composition of
601 water masses (panels g, h, i, and j) and potential changes in the characteristic
602 size of eddies (panel k). Differences between today’s reference reanalyses and
603 paleo $\times 3$ CO_2 results from a combination of factors: biases in low-resolution
604 GCMs, the chosen Eocene geography in the DeepMIP configurations, and
605 the CO_2 concentrations used in the simulations. Figure 8 quantifies these
606 differences, helping to identify their sources (main contributing factors are
607 highlighted) and assess the reliability of these changes. However the individ-
608 ual contributions of model biases, paleogeography, and CO_2 forcing are not
609 simply additive, and their combined effect may involve nonlinear interactions

610 that we are not able to quantify within the scope of this study. In Figure 8,
 611 we have also listed the models which deviate from the main results, as well
 612 as the referenced proxy studies related to these results (right column of each
 613 panel). Prior to assessing the influence of Eocene-specific features such as
 614 elevated CO₂ concentrations and paleogeography, it is crucial to identify the
 615 key biases arising from the models themselves. The pre-industrial (PI) sim-
 616 ulations from the DeepMIP ensemble constitute an essential benchmark to
 617 evaluate the impact of Eocene boundary conditions on ocean circulation. It is
 618 well established that low-resolution GCMs do not explicitly resolve mesoscale
 619 turbulence, despite its dominant contribution to the oceanic kinetic energy
 620 and its central role in lateral transports of heat, salt, and buoyancy, as well
 621 as intrinsic variability (Ferrari and Wunsch, 2009; Danabasoglu et al., 2014).
 622 This limitation leads to persistent biases in western boundary currents and
 623 frontal systems, affecting air–sea exchanges and gyre strength (Danabasoglu
 624 et al., 2014; Hewitt et al., 2017), and prevents a realistic representation of
 625 submesoscale and mixed-layer processes, often resulting in overly deep mixed
 626 layers and biased vertical exchanges that are highly sensitive to numerical
 627 dissipation (Bodner et al., 2023). Deep convection, high-latitude water-mass
 628 transformation, and overflows are likewise affected. While they do not re-
 629 solve mesoscale features due to their coarse resolution (2°), they still manage
 630 to reproduce key aspects of the present-day South Indian Ocean circulation.
 631 The PI simulations capture the large-scale horizontal circulation patterns of
 632 the SISG, including a well-positioned STF (Figure 8c), delimiting the ACC to
 633 the south and the anticyclonic South Indian Subtropical Gyre system to the
 634 north. However, certain aspects such as the magnitude of the Agulhas leak-
 635 age (Figure 8b) and the overall gyre strength (Figure 8e) exhibit substantial
 636 intermodel variability. In terms of water mass structure, the PI simulations
 637 are able to replicate the general layering and shape of temperature–salinity
 638 (Θ – S) profiles within the subtropical gyre (Figure 4). The Antarctic Inter-
 639 mediate Water (AAIW), for instance, is represented in most models as a
 640 distinct low-salinity water mass subducting beneath the Subtropical Surface
 641 Water (STSW). However, salinity and temperature biases are apparent, par-
 642 ticularly in the intermediate and deep layers, leading to a negative bias in
 643 modeled density profiles compared to WOA. Yet, some models show devi-
 644 ations in Rossby deformation radius due to differences in bathymetry and
 645 vertical stratification structure, which can affect the accurate simulation of
 646 western boundary current dynamics.

5.2. Transport changes

The barotropic transport streamfunction differs radically between the present day (GLORYS reanalysis) and the DeepMIP Eocene x3 CO₂ simulations. The ACC is absent in the paleo x1 CO₂ configuration and is replaced by a cyclonic subpolar gyre (Figure 8a), due to closed Tasman Strait and Drake Passage. The position of the zero wind stress curl indicates the boundary between the subtropical gyre and the subpolar gyre and links changes in wind stress curl and transport. Several factors can explain a weak or collapsed Agulhas leakage. R  hs et al. (2022) links the modern increase in Agulhas leakage associated with climate warming to the intensification of wind stress curl in the region. In our case, under the paleo x1 CO₂ configuration, the amplitude of the wind stress curl tends to decrease (Figure 2), which could account for a reduction in Agulhas leakage. Additionally, Verma et al. (2023), through the study of planktonic foraminifera and other proxies, show that during glacial periods, the STF shifted 7   further north, blocking the access of Indian subtropical waters to the Atlantic. In our scenario, the African continent is located 10   further south (Lotfy et al., 2017), reaching 40  S, which defines the position of the STF and effectively prevents Agulhas leakage. With the zero wind stress curl shifted northward (Figure 8d), it moves significantly closer to the African continent, which leads to a reduction in Agulhas leakage (Sijp and England, 2008). Several other proxies have been employed to investigate past changes in the Agulhas leakage. For instance, Peeters et al. (2004) relied on planktonic foraminiferal assemblages, while Mart  nez-M  ndez et al. (2010) used $\delta^{18}\text{O}$ and Mg/Ca ratios measured on *Globigerina bulloides* from the Agulhas corridor during glacial periods. However, such records are not available for the Eocene.

The position of the retroflexion is also affected by the paleo-configuration. Indeed, Africa, which intersects the STF, prevents the retroflexion from occurring south of Africa. It now takes place to the east of the African coast. A displacement of the retroflexion could have substantial consequences for both oceanic and climatic conditions in the region, one of which might be its influence on the regional hydrological cycle. Njouodo et al. (2021) show that the Agulhas Current system plays a very important role in precipitation in the region. This result is also evident in Figure 4 of Williams et al. (2022), which compares precipitation patterns from the DeepMIP early Eocene simulations. For both paleo x1 CO₂ and paleo x3 CO₂ scenarios, southern Africa experiences reduced precipitation, whereas the ocean to the east of the continent where the retroflexion is now located shows increased precipitation

685 (except in the GFDL model).

686 The primary factor explaining the reduction in transport (Figure 8e) is
687 the paleogeographic change in the Eocene, but the increase in CO_2 also in-
688 duces a further decrease in the gyre transport from an average of 48.6 Sv in
689 the paleo x1 CO_2 simulations to 39.9 Sv in the paleo x3 CO_2 simulations.
690 This corresponds to a reduction comparable to the decrease in the inten-
691 sity of the wind stress curl of the westerlies, suggesting that the subtropical
692 gyre circulation during the Eocene was primarily Sverdrupian i.e., driven by
693 the balance between wind stress curl and the meridional transport of wa-
694 ter as described by Sverdrup theory (Sverdrup, 1947), a feature that is also
695 observed in the present day (Penven et al., 2021). Kirtland Turner et al.
696 (2024) demonstrate, using benthic foraminiferal $\delta^{13}\text{C}$ and $\delta^{18}\text{O}$ records along
697 with an intermediate-complexity model, that overturning circulation in the
698 Southern Hemisphere declined during the Early Eocene.

699 The Eocene x3 CO_2 simulation shows that the STF is located 4.4° fur-
700 ther south compared to the present-day position. As with other features,
701 the Eocene geography is the primary driver of this shift (see Figure 8), while
702 CO_2 only induces a minor southward displacement of 0.2° . Studying proxies
703 including dinoflagellate cyst biogeography and benthic foraminiferal clumped
704 isotopes from offshore Tasmania, Hou et al. (2023) suggest that during the
705 Eocene, the STF was located even farther south (53°S) before migrating
706 northward between 14 Ma and 7 Ma to its present-day position around 40°S .
707 This suggests that models may underestimate the southward shift. De Boer
708 et al. (2013) show that this southward shift of the STF is not incompati-
709 ble with the northward migration of the westerlies (51°S to 47.3°S) and that
710 the STF position is primarily controlled by topography, explaining why geo-
711 graphic changes remain the dominant factor in its displacement.

712 5.3. *Water masses changes*

713 We have observed that water masses are also profoundly affected in these
714 different simulations. The Eocene is known for its warmer oceans compared
715 to the present day. In our study region and for our simulations, this warming
716 is particularly striking at depth. The average bottom temperature over our
717 characteristic gyre region (50°E – 100°E , 20°S – 40°S) increases from 0.9°C in
718 WOA to 10.4°C in the Eocene x3 CO_2 simulations (Figure 8g). Our find-
719 ings are consistent with those of Zhang et al. (2020), who estimated a similar
720 deep-water temperature increase of this magnitude using Mg/Ca ratios. This

721 is also in line with Agterhuis et al. (2022) which measured stable and agglom-
722 erated isotopes by combining carbonate clumped isotope thermometer (Δ_{47})
723 measurements from different benthic species, obtaining temperatures during
724 hyperthermal events ranging from 13°C to 17°C. More broadly, compilations
725 of long-term CO₂ reconstructions demonstrate that such elevated greenhouse
726 gas levels are sufficient to maintain substantially warmer deep-ocean condi-
727 tions over geological timescales (Foster et al., 2017). While this temperature
728 change is primarily driven by the increase in CO₂, geography also plays a
729 significant role, accounting for 27% of the warming, in contrast to the PI
730 simulations, which yield bottom temperatures identical to WOA. This effect
731 could be attributed to the closure of major oceanic gateways, which has been
732 shown to strongly modulate ocean circulation and deep-water properties in
733 Eocene boundary condition experiments (Toumoulin et al., 2020).

734 Keeping in mind the sea surface/subsurface salinity model inter-variability,
735 the maximum salinity (around 200 m depth) in the sections of our study re-
736 gion (Figure 6b and Figure 8h) increases by 0.5 PSU between WOA and the
737 Eocene x3 CO₂ simulations. Once again, geographical changes appear to be
738 the primary driver of this result. A saltier Indian Ocean during the Eocene
739 is also supported by proxy data. At Ocean Drilling Program (ODP) Site
740 1209 in the Pacific, which provides insights into broader tropical ocean pat-
741 terns, including the Indian Ocean, an increase in Sea Surface Salinity (SSS)
742 during the Paleocene-Eocene Thermal Maximum was found (Harper et al.,
743 2018; Carmichael et al., 2016). These studies, combining model comparisons
744 and proxy data, demonstrate a global intensification of the hydrological cy-
745 cle, with significantly higher evaporation-precipitation (E-P) values between
746 0°S and 40°S. It is noteworthy that orbital cycles, which remain unchanged
747 in this case, also played a significant role in shaping hydrological cycles, as
748 highlighted by Briard (2020). Further sensitivity tests would be necessary to
749 fully understand the exact role of geography in the observed salinity increase.

750 The temperature gradient between low and high latitudes is significantly
751 affected in the Eocene x3 CO₂ configurations (Figure 8j) compared to the
752 present day, decreasing from 27.8°C in WOA to 21°C in x3 CO₂. Numer-
753 ous proxy sources, including leaf assemblages, floristic composition, as well
754 as $\delta^{18}O$ isotopes in planktonic and benthic foraminifera, support this result
755 (Willard et al., 2019; West et al., 2020; Eichenseer and Jones, 2024; Gaskell
756 et al., 2022). This result, known as polar amplification, is also observed in
757 current climate change, where the Arctic is warming on average four times
758 faster than the global climate, and Antarctica three times faster (Rantanen

et al., 2022; Clem et al., 2020). In the past, this phenomenon was further amplified by paleogeography. Indeed, with closed oceanic gateways, the meridional heat transport was significantly enhanced (Zhang et al., 2020).

All DeepMIP models are able to generate AAIW in the PI simulations characterized by a freshwater peak around the isopycnal 1036 kg m^{-3} in Figure 4. They do not exhibit any evidence of AAIW formation in the paleo x1 CO_2 and paleo x3 CO_2 simulations, except for GFDL. The strong westerlies driving the Antarctic Circumpolar Current (ACC) generate equatorward Ekman transport, which plays a crucial role in the formation of AAIW. This process pushes surface waters away from Antarctica, contributing to the Antarctic Divergence region and inducing the upwelling of deep water, which subsequently subducts beneath the subtropical surface water, forming AAIW (Sallée et al., 2010). Around 42 Ma, a significant shift occurred in the Indian Ocean, leading to a divergence in $\delta^{18}\text{O}$ values (Holmström, 2014) between its northern and southern regions. This divergence suggests the emergence of a thermal segregation of intermediate water masses, resulting in two distinct water bodies: the cool, high-latitude AAIW and a warm, low-latitude water mass known as Tethyan-Indian Saline Water (TISW). This phenomenon is believed to be linked to the formation of the ACC and the cooling of high-latitude temperatures, which allowed intermediate waters to sink more extensively than before, forming the cool water mass on a larger scale. Further evidence of this thermal segregation is provided by changes in the distribution of planktonic and benthic foraminifera, as well as a marked retreat of warm water calcareous nannoplankton in the southern Indian Ocean (Villa et al., 2014).

5.4. *Wind changes*

DeepMIP simulations under Eocene conditions (CO_2 concentrations and paleogeography), represented in the paleo x3 CO_2 and paleo x1 CO_2 configurations, indicate a northward shift and weakening of the westerly winds compared to the modern situation (ERA5 reanalysis). This shift differs from the findings of Chen et al. (2024), who suggested a more southerly position of the westerlies during the Eocene, based on neodymium and hafnium isotopic ratios and clay mineralogy extracted from the U1514 site in the Mentelle Basin. While the weakening wind stress curl is not directly measured in proxies, Eichenseer and Jones (2024) investigated this aspect using geochemical proxies combined with ecological data from mangrove communities and

coral reefs. Their study estimated a flattened latitudinal sea surface temperature (SST) gradient of 13.3°C. In contrast, DeepMIP simulations give an average SST gradient of 21°C (Figure 8j), which remains lower than the present-day gradient of 27.8°C observed in WOA. This reduction in the SST gradient, when applied to the thermal wind balance, leads to a weakened westerly wind system. The lower meridional temperature gradient results in reduced Sea Level Pressure (SLP) over Antarctica compared to stronger mid-latitude SLP, thereby driving a poleward shift of the westerlies. This mechanism is currently observed in global warming scenarios and is associated with the positive phase of the Southern Annular Mode (SAM) (Marshall, 2003). As indicated in Figure 8, the most significant shift (2.6°) in the position and intensity of the westerlies occurs in the ERA5/PI comparison. This suggests that the primary driver of the westerly wind shift is the PI models biases. Therefore, this result should be interpreted with caution. Nevertheless, we still obtain a northward shift of the westerlies between PI and paleo x1/x3 CO₂. Similarly, Abhik et al. (2024) found that monsoonal changes were predominantly controlled by paleogeographic factors, with atmospheric CO₂ playing a minor role. Our results also indicate a stronger influence of paleogeography compared to atmospheric CO₂ concentrations. These important PI model biases, with strong inter-model variability, can be attributed to various factors, including the parameterization of atmospheric turbulent fluxes, heat and moisture exchanges, coupling realism, model resolution, cloud parameterization, and surface albedo representation. Additionally, pre-industrial simulations are performed with lower atmospheric CO₂ concentrations than ERA5, which may contribute to the observed shift. Indeed, ongoing global warming is causing a poleward shift and intensification of the Southern Hemisphere westerlies (Tim et al., 2019; Perren et al., 2020).

5.5. High resolution paleoceanographic perspectives

A smaller Rossby deformation radius, as observed in the paleoclimate simulations for the Eocene (paleo x1 CO₂ and paleo x3 CO₂ in Figure 7b), could significantly impact ocean dynamics in the region. The Rossby deformation radius sets the size and propagation speed of Rossby waves and eddies. Outside the equatorial band, Chelton et al. (1998) defines the Rossby deformation radius (λ_m) associated with mode- m as $\lambda_m = \frac{c_m}{f}$, where c_m is the phase speed of gravity waves for mode- m and f is the coriolis parameter. The magnitude of this radius is critical for many oceanic processes.

Chelton et al. (2011) demonstrated that smaller eddies (corresponding to a smaller Rossby deformation radius) dissipate more rapidly, leading to enhanced energy dissipation via friction and turbulent mixing. When the Rossby deformation radius is smaller, the propagation speed of first-mode baroclinic Rossby waves decreases. This implies that wind anomalies take longer to influence ocean circulation because pycnocline adjustment through these waves becomes slower. Hutchinson et al. (2018) showed that the seasonal phasing of the Agulhas Current is strongly dependent on this propagation speed: when baroclinic waves propagate more slowly, the current’s response to atmospheric forcing is delayed. They also showed that signals from distant winds tend to dissipate before reaching the western boundary of the basin, mainly due to destructive interference between local wind anomalies and those propagating from the east. In a context of a smaller Rossby deformation radius, the current’s response would thus be even more confined to local or near-field forcing, as long-distance communication via Rossby waves would be further inhibited. This restricts the influence of remote winds and strengthens the direct role of regional winds in modulating western boundary circulation variability.

Theiss (2004) showed that there exists a critical latitude defined by the ratio r between the Rossby deformation radius and the Rhines scale: $L_R = \sqrt{\frac{U}{\beta}}$, where U is the characteristic turbulent velocity and β is the meridional gradient of the Coriolis parameter. When $r > 1$ (at low latitudes), Rossby waves dominate and induce anisotropy by organizing the flow into alternating zonal jets. For $r < 1$, Rossby waves are less effective at structuring the flow, resulting in isotropic turbulence dominated by vortices. In practice, the critical latitude (where $r = 1$) is estimated to lie between 10° and 20° latitude (Halo et al., 2023). A smaller Rossby deformation radius would shift the condition $r = 1$ to lower latitudes. This means that zonal structures, such as alternating jets, become confined to lower latitudes, while isotropic, vortex-dominated turbulence extends closer to the equator.

Western boundary currents balance the transport of subtropical gyres in a boundary layer often associated with topography (Hughes, 2000; Gula et al., 2015). Nevertheless, frontal arguments have been put forward to link their width with the Rossby deformation radius (Saenko, 2006). As a consequence a smaller Rossby deformation radius should make western boundary currents narrower, more intense (to maintain the necessary transport) and more stable (Saenko, 2006). Chelton et al. (1998) also showed that mesoscale variability

868 and the width of currents are directly linked to the geographical distribution
869 of the Rossby deformation radius.

870 For modeling the mesoscale structures of the Agulhas Current in a pa-
871 leo context, it is useful to compare Rossby deformation radius south of the
872 Agulhas Bank. From GLORYS reanalysis, Rossby deformation radius is es-
873 timated around 51 km, whereas in the Eocene x3 CO₂ simulations, estimates
874 are around 22 km (Figure 8k). This difference is particularly significant,
875 because Africa was about 10° further south during the Eocene (Lotfy et al.,
876 2017), and Rossby deformation radius decreases with latitude. According
877 to Soufflet et al. (2016), simulations with model based on high order nu-
878 merics require at least a resolution seven times finer than the scale of the
879 phenomenon to be properly resolved. Thus, for a Rossby deformation ra-
880 dius of 22 km (assuming structures span approximately $2R_0$), a resolution of
881 around 6 km is needed to adequately resolve mesoscale features in Eocene
882 simulations. By contrast, for the present-day Rossby deformation radius of
883 51 km, a resolution of about 14 km is required.

884 Importantly, these requirements are far finer than the ocean grid spacing
885 used in the DeepMIP coupled simulations analyzed here. Consequently, the
886 present DeepMIP simulations cannot resolve mesoscale vortices, ring shed-
887 ding, or the associated kinetic energy cascade in the Agulhas system, and
888 they are therefore likely to significantly underestimate the eddy kinetic en-
889 ergy as well as the associated temporal and spatial variability during the
890 Eocene. The Rossby deformation radius-based arguments presented above
891 should therefore be regarded as a physically grounded indication of how the
892 unresolved mesoscale regime may differ under Eocene boundary conditions,
893 since the dynamical consequences of a reduced Rossby deformation radius
894 cannot be captured by the current low-resolution models.

895 In order to resolve the mesoscale processes of the Agulhas Current system,
896 Le Hir et al. (2024) conducted a sensitivity test of ocean circulation in the
897 southwestern Indian Ocean to the closure of the Mozambique Channel, which
898 may have occurred intermittently during the Cenozoic (Masters et al., 2021;
899 Aslanian et al., 2023), using a model with horizontal resolution of 1/12° (i.e.
900 about 9 km). However, Le Hir et al. (2024) remains an idealized study and
901 cannot be considered as a paleoclimatic experiment, as large-scale boundary
902 conditions and atmospheric forcing were unchanged. Results we obtained
903 here will serve as a preliminary step in identifying robust patterns and the
904 biases inherent in GCMs in their representation of the large-scale SISG in
905 the Eocene. Le Hir et al. (2024) provides insights into the tools to apply the

906 sensitivity test of ocean circulation in the southwestern Indian Ocean to the
 907 closure of the Mozambique Channel in a paleoclimatic context. Given Rossby
 908 deformation radius relative to Africa’s position, eddy-resolving regional sim-
 909 ulations in the Eocene configuration would require a finer resolution than for
 910 the present to resolve explicitly the mesoscale processes.

911 6. Conclusion

912 The South Indian Subtropical Gyre during the Early Eocene was substan-
 913 tially different from its modern configuration. Simulations from the DeepMIP
 914 project show a weaker, restructured gyre extending into the Pacific, bordered
 915 in the south by a subpolar gyre replacing today’s Antarctic Circumpolar Cur-
 916 rent and at North by a tropical gyre, disconnected from the Atlantic due to
 917 the collapse of the Agulhas leakage, and a poleward-shifted STF. Despite lim-
 918 itations of these low-resolution General Circulation Models, DeepMIP models
 919 consistently reproduce key features mainly supported by proxy data studies,
 920 such as warmer deep waters, a reduced meridional temperature gradient, the
 921 disappearance of Antarctic Intermediate Water (AAIW), and altered wind
 922 patterns. The dominant influence of Eocene paleogeography, especially the
 923 closure of ocean gateways and the southward shift of the African continent,
 924 emerges as the main driver of these changes, more than CO₂ influence. Over-
 925 all, the results provide robust insights into past ocean dynamics and lay the
 926 groundwork for future studies that aim to refine our understanding of the
 927 ocean circulation under past climate conditions.

928 Appendix A: Method to compute the barotropic streamfunction

929 ψ represents the depth-integrated horizontal mean flow in the ocean and
 930 reflects the position, strength, and connections of the gyre with other basins.
 931 It represents the mean transport associated with the large-scale circulation.
 932 The streamfunction can be derived from the Helmholtz decomposition, which
 933 states that any vector field can be decomposed into an irrotational (divergent)
 934 component and a non-divergent (rotational) component:

$$\mathbf{u} = \nabla \times \psi + \nabla \phi,$$

935 where $\nabla \times \psi$ is the rotational (non-divergent) component and $\nabla \phi$ is the
 936 irrotational (divergent) component. Taking the curl of the barotropic velocity
 937 field yields:

$$\Delta \psi = \nabla \times \mathbf{u} = \zeta$$

938 where ζ is the vorticity of the vertically integrated currents, $\zeta = \partial_x V - \partial_y U$.
 939 The following method is applied to invert the Laplacian and compute ψ :

- 940 1. Verify that the fluxes of U and V are globally conserved over the
 941 selected domain (here encompassing the SISG). If not, u and v are
 942 adjusted to enforce global mass flux conservation along the domain
 943 boundaries.
- 944 2. Compute ψ at the boundaries from U and V . Propagate ψ across the
 945 land mask, ensuring that ψ remains constant along solid boundaries
 946 (no flux through land).
- 947 3. Identify and label connected regions (e.g., islands) in a binary mask
 948 by assigning a unique integer to each contiguous group of land masks
 949 elements based on specified connectivity.
- 950 4. Since the problem is linear, the solution can be expressed following
 951 Colin de Verdière and Ollitrault (2016) as:

$$\psi = \psi_0 + \sum_{k=1}^N \mu_k \psi_k$$

952 where:

- 953 • ψ_0 is the circulation on the entire domain, it is the particular
 954 solution of the Poisson equation $\Delta\psi_0 = \nabla \times \mathbf{u}$, with zero Dirichlet
 955 boundary conditions on the external boundary and on all islands
 956 ($\psi_0 = 0$).
- 957 • ψ_k is the circulation associated with island k , it is the homogeneous
 958 solution satisfying $\Delta\psi_k = 0$, with $\psi_k = 1$ on island k and $\psi_k = 0$
 959 on all other boundaries and islands.

960 ψ_0 and ψ_k are determined by applying a successive over-relaxation
 961 method (Yang and Matthias, 2007) to invert the Laplacian. The coef-
 962 ficients μ_k are determined to ensure that the circulation around each
 963 island (defined by the contour ∂I_k) matches the circulation derived
 964 from the barotropic transport field $\mathbf{u}$:

$$\underbrace{\oint_{\partial I_k} \mathbf{u} \cdot \mathbf{n} ds}_{\text{Circulation around island } k} = \underbrace{\oint_{\partial I_k} \partial_n \psi_0 ds}_{\text{Contribution from } \psi_0} + \sum_{j=1}^N \mu_j \underbrace{\oint_{\partial I_k} \partial_n \psi_j ds}_{\text{Contribution of each islands}}$$

965 This forms a linear system, for $k = 1, \dots, N$:

$$\sum_{j=1}^N A_{kj} \mu_j = b_k$$

966 with:

$$A_{kj} = \oint_{\partial I_k} \partial_n \psi_j ds, \quad b_k = \oint_{\partial I_k} \mathbf{u} \cdot ds - \oint_{\partial I_k} \partial_n \psi_0 ds$$

967 Solving this system yields the coefficients μ_k , from which the complete
 968 streamfunction ψ can be reconstructed to satisfy both the vorticity
 969 constraint and the correct circulation around each island.

970 This method ensures that the streamfunction accurately represents the cir-
 971 culation around each island in the domain, as imposed by the barotropic
 972 transport field.

973 With the exception of IPSL-CM5A2, the models provide an output vari-
 974 able for the barotropic streamfunction (“msftbarot”), as described in the
 975 CMIP6 documentation (Griffies et al., 2014). This variable is computed by
 976 vertically integrating the zonal mass transport (ρu) over the entire water
 977 column and then meridionally integrating this transport from Antarctica.
 978 The method we propose, ensures a dynamically consistent circulation by ex-
 979 plicitly enforcing mass conservation at the boundaries of the domain. This
 980 distinction is particularly important in our study, as the u and v velocity
 981 fields are interpolated onto a common, finer grid across all simulations in
 982 order to compute and compare key metrics of Agulhas leakage. As shown in
 983 Figure A1, the choice of method has little impact on the pre-industrial sim-
 984 ulations of MIROC. In contrast, for HadCM3B, the Antarctic Circumpolar
 985 Current transport is reduced, and the gyre recirculation is approximately 15
 986 Sv weaker. These differences may have several sources. First, the archived u
 987 and v fields are temporally averaged, whereas “msftbarot” is diagnosed di-
 988 rectly within the model. Second, the interpolation of the velocity fields onto a
 989 common grid may introduce small local inconsistencies in the vertically inte-
 990 grated transport. In a cumulative integration, such small differences can ac-
 991 cumulate with distance from the southern boundary, leading to discrepancies
 992 in the reconstructed streamfunction even when the underlying model solution
 993 itself is mass-conserving. In this context, the Laplacian-inversion/Helmholtz
 994 approach is used not because the model transport is not mass-conserving,

995 but because it provides a robust and numerically controlled diagnostic of
 996 the barotropic circulation that is less sensitive to the direction of integration
 997 and better suited for intercomparison across models with different grids and
 998 geometries.

999 **Appendix B: Method to compute the Agulhas leakage**

1000 Because of the coarse resolution of the paleoclimate simulations, the Ag-
 1001 ulhas Leakage cannot be associated with eddies, allowing the use of the mean
 1002 transport function ψ to evaluate its intensity. This is achieved by specifically
 1003 separating the streamlines ending in the Atlantic Ocean (i.e. defining the
 1004 Leakage) to the one participating to the Agulhas retroflection and to the
 1005 Antarctic Circumpolar Current. Since we have defined the value of ψ to
 1006 be zero over the African continent, this is practically achieved by selecting
 1007 and summing contours spaced at 0.1 Sv that satisfy the following conditions
 1008 (summarized in Figure B1):

- 1009 • **STF**: crosses both the Good Hope Line and the 40°E section north
 1010 of the STF. This condition excludes contours associated with ACC
 1011 counter-currents.
- 1012 • **GHL**: connects the Good Hope Line (GHL) at 0°E, as defined by
 1013 Gladyshev et al. (2008). This condition ensures that the contour enters
 1014 the Atlantic Ocean and eliminates local recirculations associated with
 1015 the Agulhas retroflection and the Agulhas Subgyre when computing
 1016 the leakage.
- 1017 • **40°E**: connects to the meridional section at 40°E, as defined by Stramma
 1018 and Lutjeharms (1997). This condition ensures that the contour origi-
 1019 nates from the Indian Ocean.
- 1020 • **East → West**: this condition is satisfied if the contour connects first
 1021 to 40°E and then to the GHL. It confirms that the barotropic flow along
 1022 the contour goes from the Indian Ocean to the Atlantic.

1023 By applying this method to the annually-averaged GLORYS horizontal
 1024 transport, we obtain 13.1 Sv, which is close to the value obtained with the
 1025 Lagrangian method applied to observations (Richardson, 2007) and from
 1026 eddy resolving simulations (Schmidt et al., 2021).

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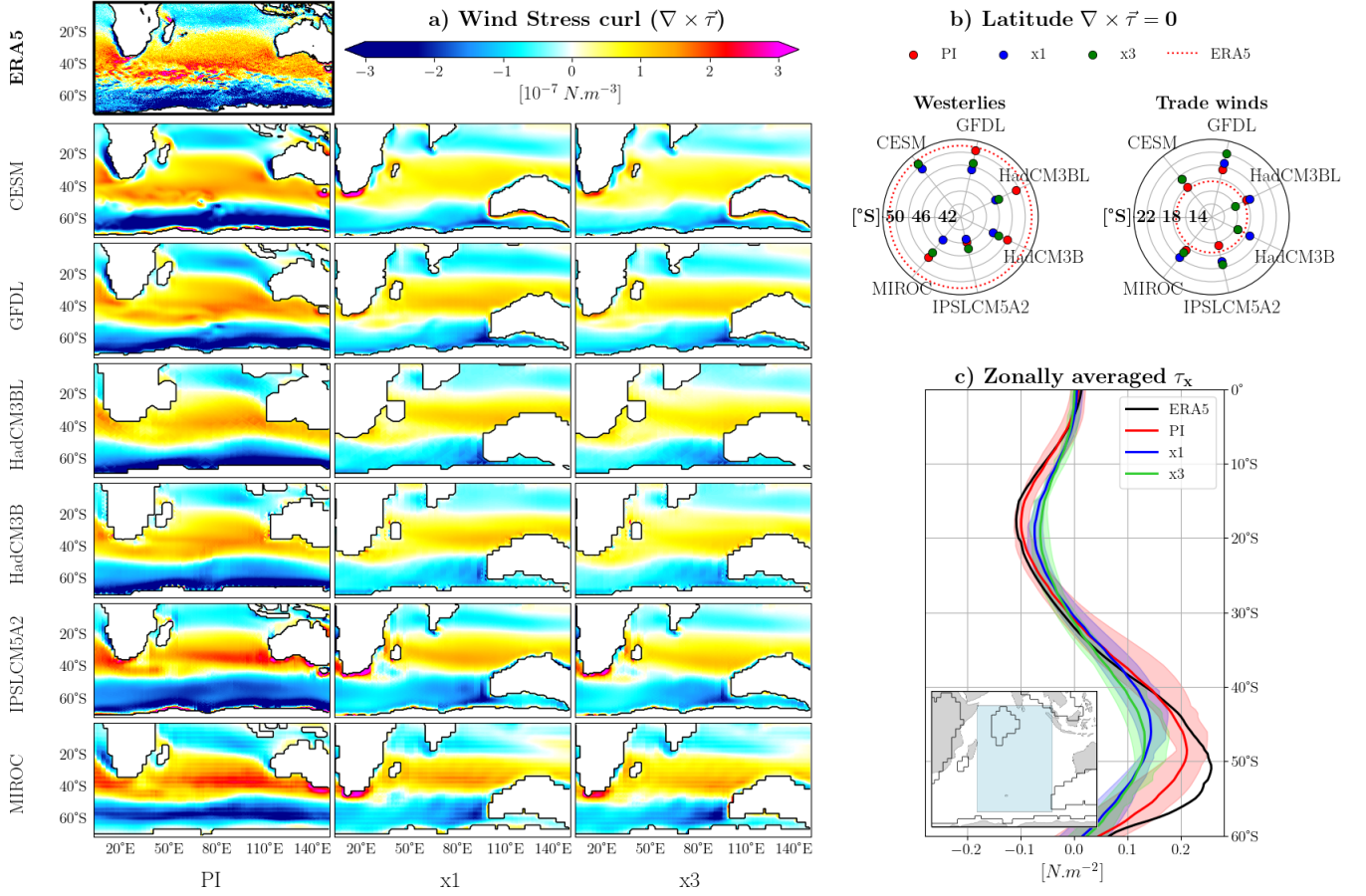


Figure 2: a) Annual-mean wind stress curl [$10^{-7} N.m^{-3}$] for the PI, paleo x1 CO₂, and paleo x3 CO₂ configurations of six DeepMIP model outputs, as well as the 2000–2009 average from the ERA5 atmospheric reanalysis. Blue/red values indicate anticyclonic/cyclonic (clockwise/anticlockwise in the Southern Hemisphere) wind forcing. b) Radial scatter plot, showing the deviation of the maximum westerlies and trade winds in degrees, defined as the mean latitude of the zero wind stress curl between 50°E and 100°E (blue shaded area on the map). c) Annual-mean zonal wind stress (τ_x , positive eastward), zonally-averaged over 50°E–100°E, for the ERA5 reference (black), averaged over the PI configurations (red), paleo x1 CO₂ configurations (blue), and paleo x3 CO₂ configurations (green). The associated shaded areas indicate the inter-model spread.

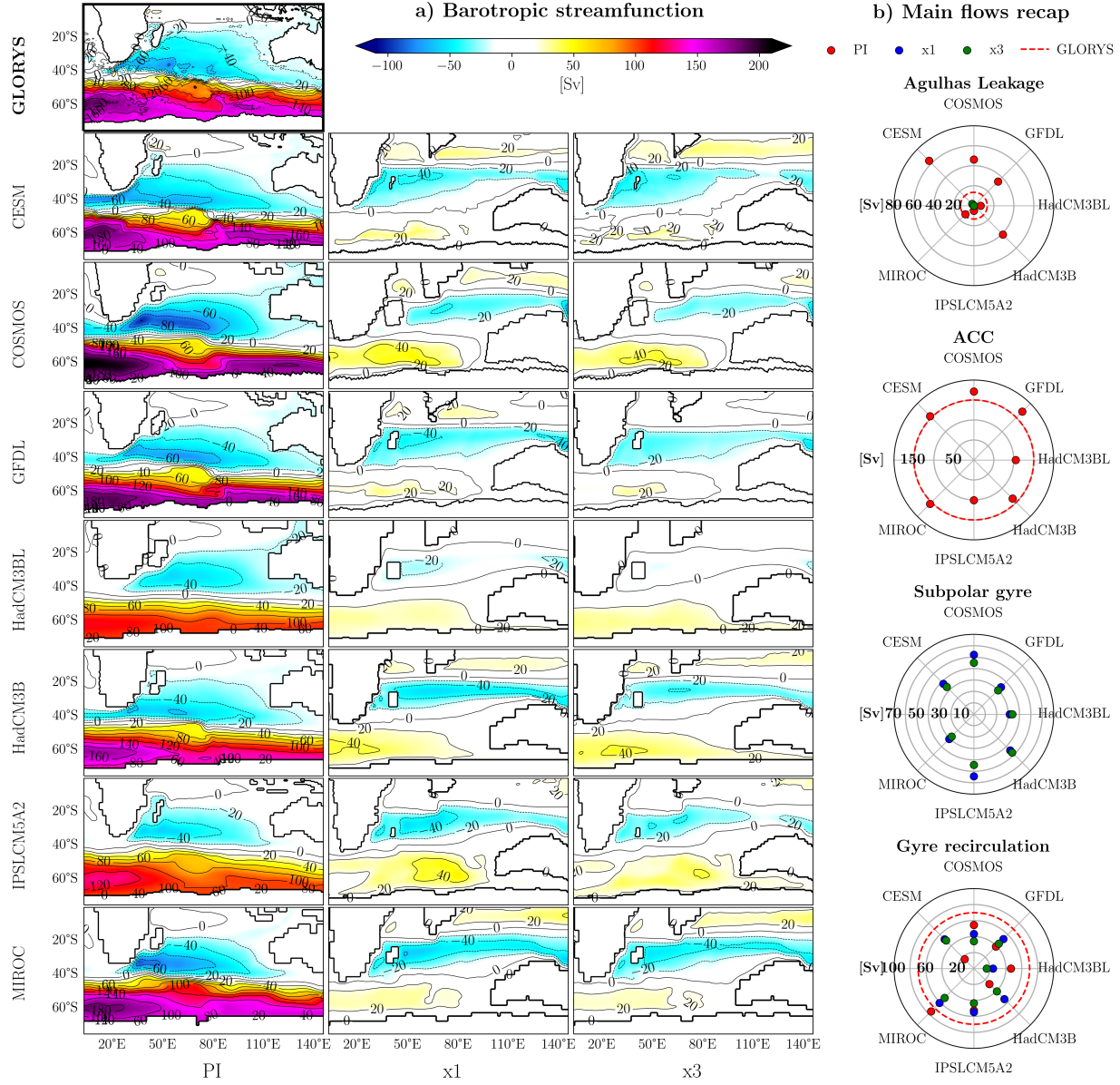


Figure 3: a) Mean barotropic streamfunction [Sv] for the PI, paleo $\times 1$ CO₂, and paleo $\times 3$ CO₂ configurations from seven DeepMIP model outputs, and for the GLORYS reanalysis averaged over 2000–2009. b) Radial scatter plots of the main flows (Aguilhas leakage, ACC, Subpolar Gyre, and Retroflexion) for GLORYS and each DeepMIP simulation. Note that only PI models are shown in the ACC scatter plot, as the ACC is absent in the paleo $\times 1$ CO₂ and paleo $\times 3$ CO₂ configurations. Similarly, the PI configuration is not shown in the Subpolar Gyre scatter plot, as this gyre exists only in the paleo $\times 1$ CO₂ and paleo $\times 3$ CO₂ configurations.

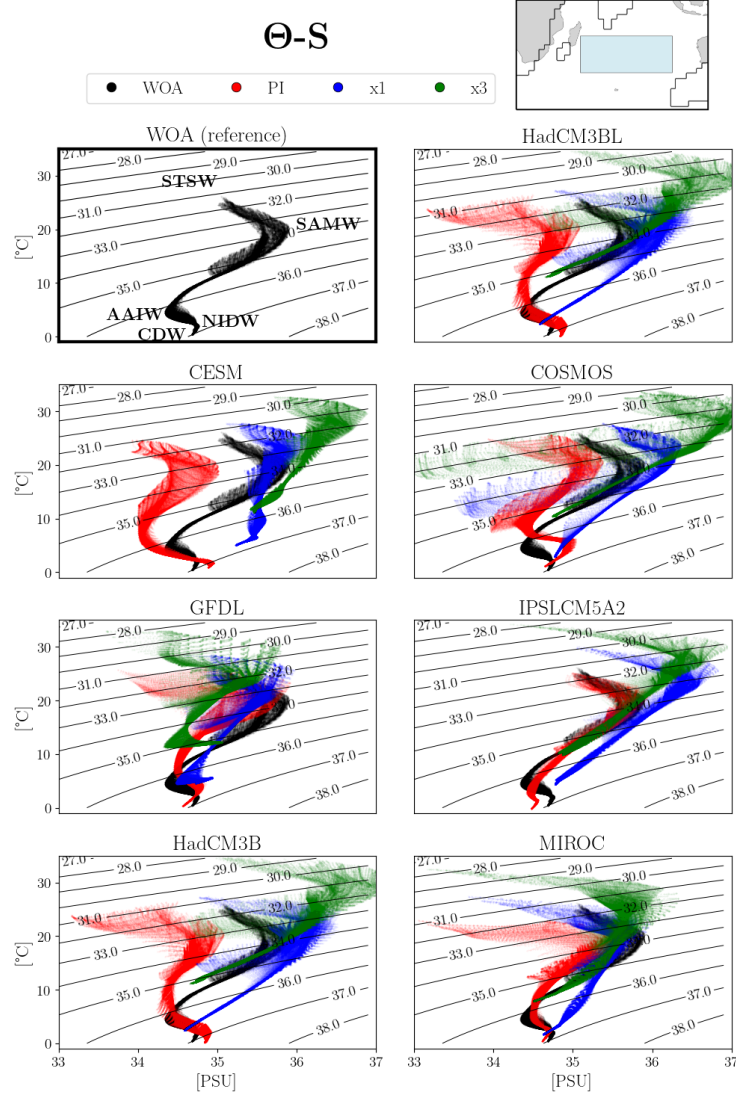


Figure 4: Θ - S diagrams for the central subtropical gyre (20°S – 40°S , 50°E – 100°E ; blue shaded area on the map) for the reference World Ocean Atlas (shown in black in each subplot) and 7 models under the 3 configurations. Contours indicate potential density anomaly (kg m^{-3}). Main water masses of the area are represented on the top left reference panel: STSW (SubTropical Surface Water), SAMW (South Atlantic Mode Water), AAIW (Antarctic Intermediate Water), NIDW (North Indian Deep Water) and CDW (Circumpolar Deep Water).

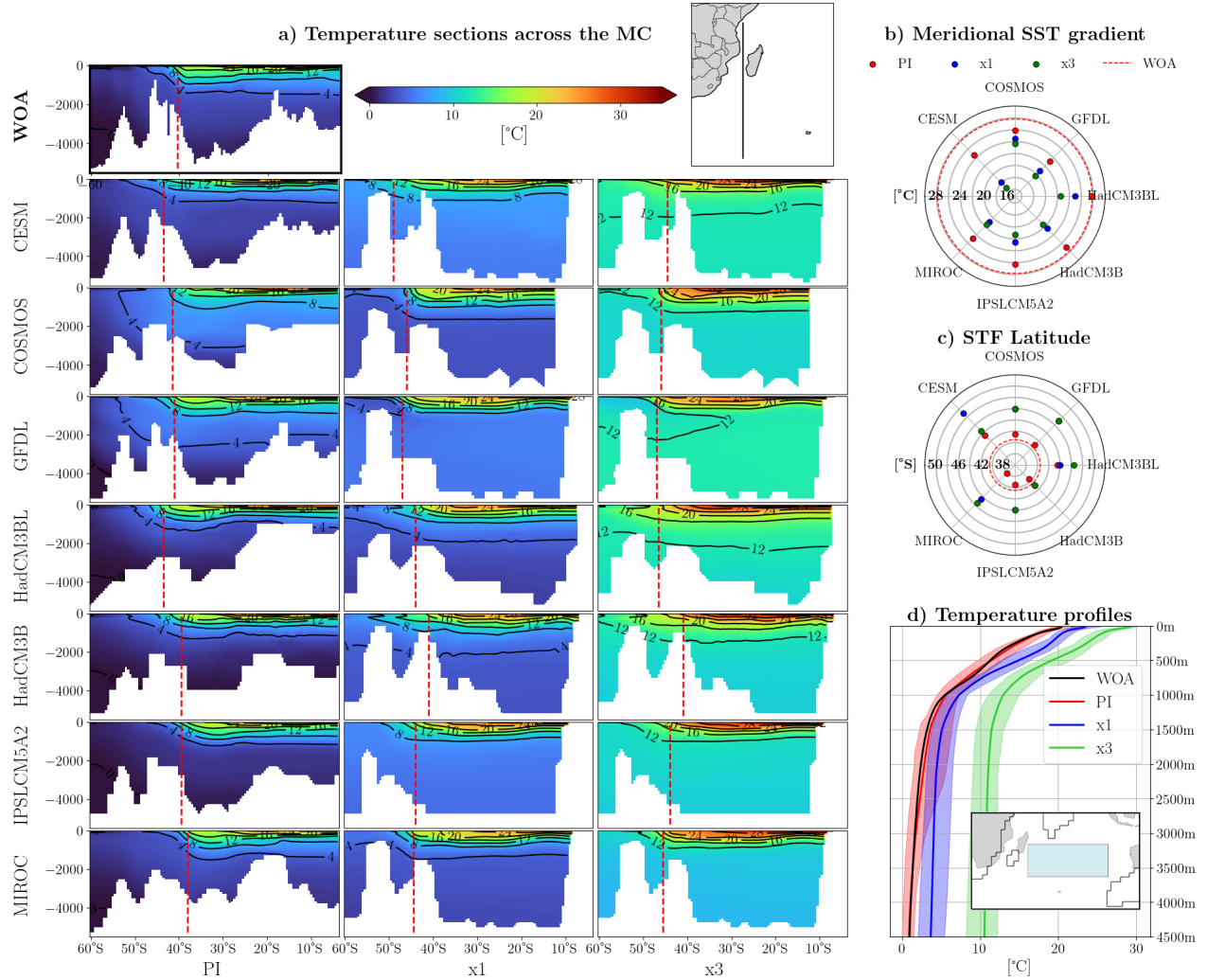


Figure 5: a) Temperature meridional section in the middle of the Mozambique Channel (42°E for the present configurations and 37°E for the paleo configurations; see map) for the PI, paleo x1 CO₂, and paleo x3 CO₂ configurations of 7 DeepMIP models and WOA. The vertical red line represents the location of the maximum meridional SST gradient. Blue/red indicate cold/warm water. b) Radial scatter plot representing maximum SST meridional gradients for each DeepMIP models. c) Radial scatter plot representing STF latitude for each DeepMIP models. d) Vertical profiles of temperature averaged over 20°S–40°S and 50°E–100°E (blue shaded area on the map) averaged over the PI configurations (red), paleo x1 CO₂ configurations (blue), and paleo x3 CO₂ configurations (green). The associated shaded areas indicate the inter-model spread.

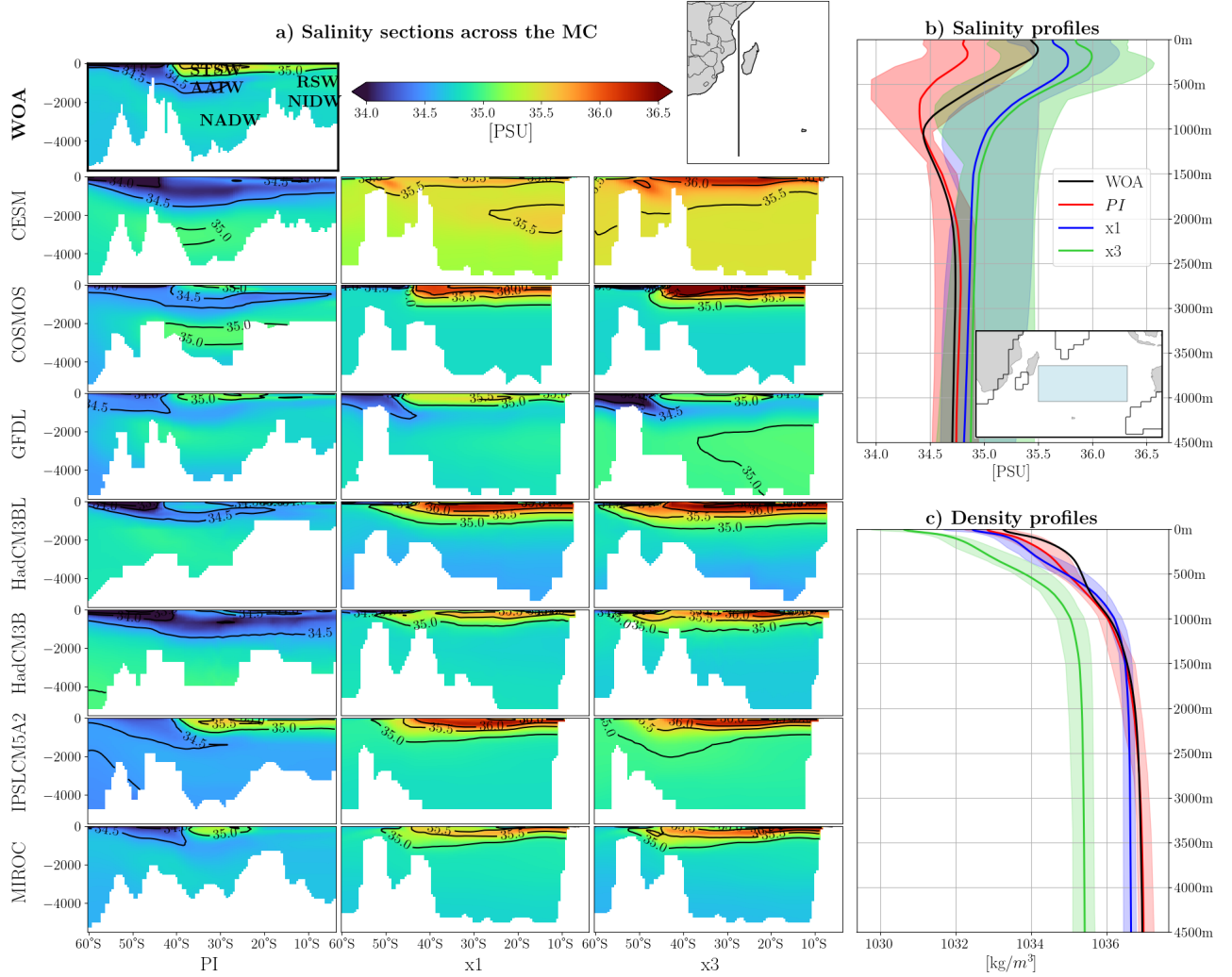


Figure 6: a) Meridional salinity sections in the middle of the Mozambique Channel (42°E for the present configurations and 37°E for the paleo configurations; see map) for the three DeepMIP configurations and WOA (top-left). Blue/red indicate fresh/salty water. Water masses are labeled in the WOA section. b) Salinity and c) density vertical profiles averaged over 20°S–40°S and 50°E–100°E (blue shaded area on the map) averaged over the PI configurations (red), paleo x1 CO₂ configurations (blue), and paleo x3 CO₂ configurations (green). The associated shaded areas indicate the inter-model spread.

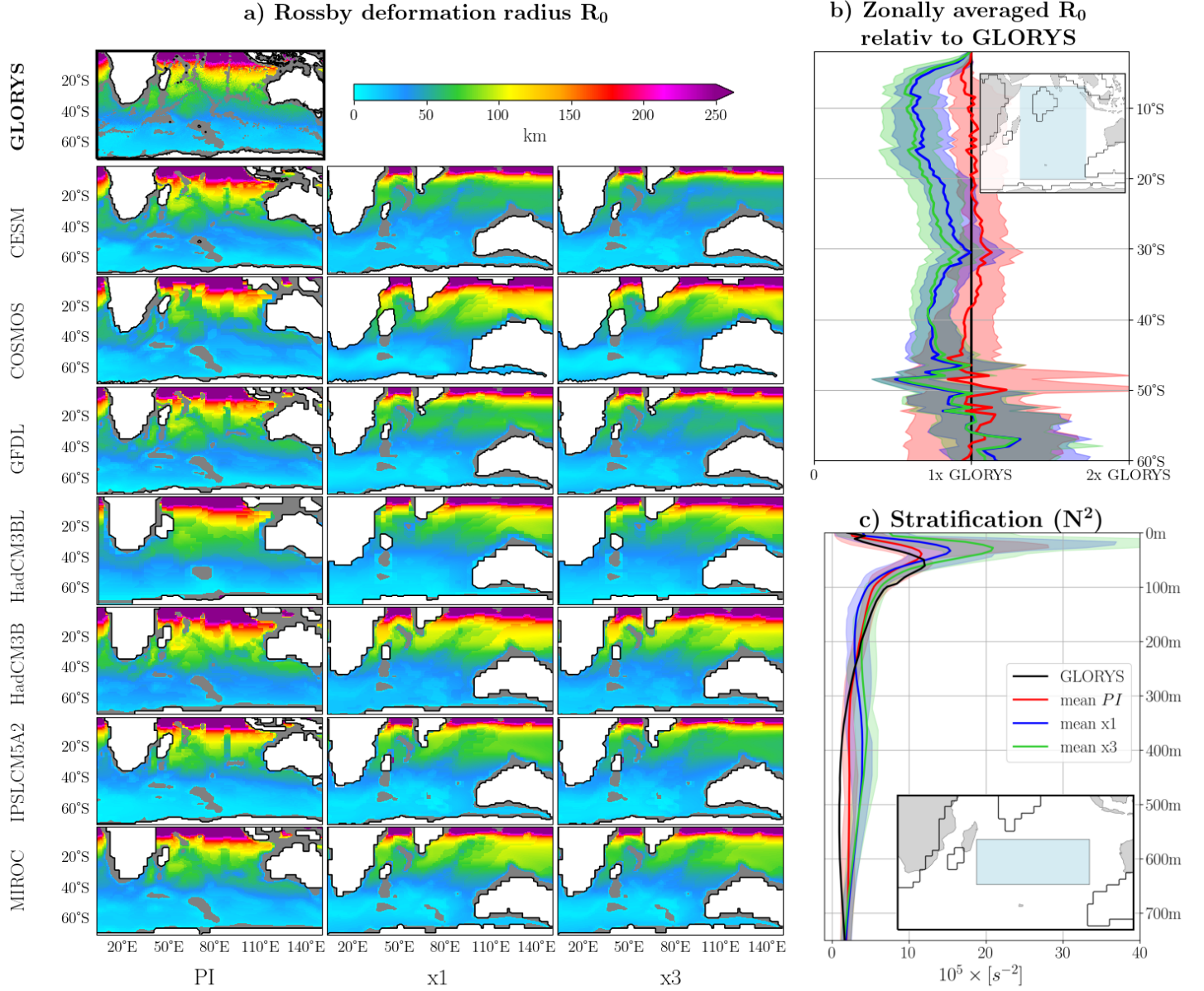


Figure 7: a) Rossby deformation radius (R_0) for the configuration PI, x1 and x3 of the DeepMIP models outputs and for GLORYS. Regions shallower than 2500 m are shaded in gray. b) Zonally-averaged Rossby deformation radius over 50°E-100°E (see map) averaged over the PI configurations (red), paleo x1 CO₂ configurations (blue), and paleo x3 CO₂ configurations (green). The associated shaded areas indicate the inter-model spread. c) Brunt-Vaisala frequency (N^2) profiles averaged in the 20°S-40°S, 50°E-100°E area (see map) averaged over the PI configurations (red), paleo x1 CO₂ configurations (blue), and paleo x3 CO₂ configurations (green). The associated shaded areas indicate the inter-model spread.

Main SISG changes found in the 56Ma DeepMIP simulations

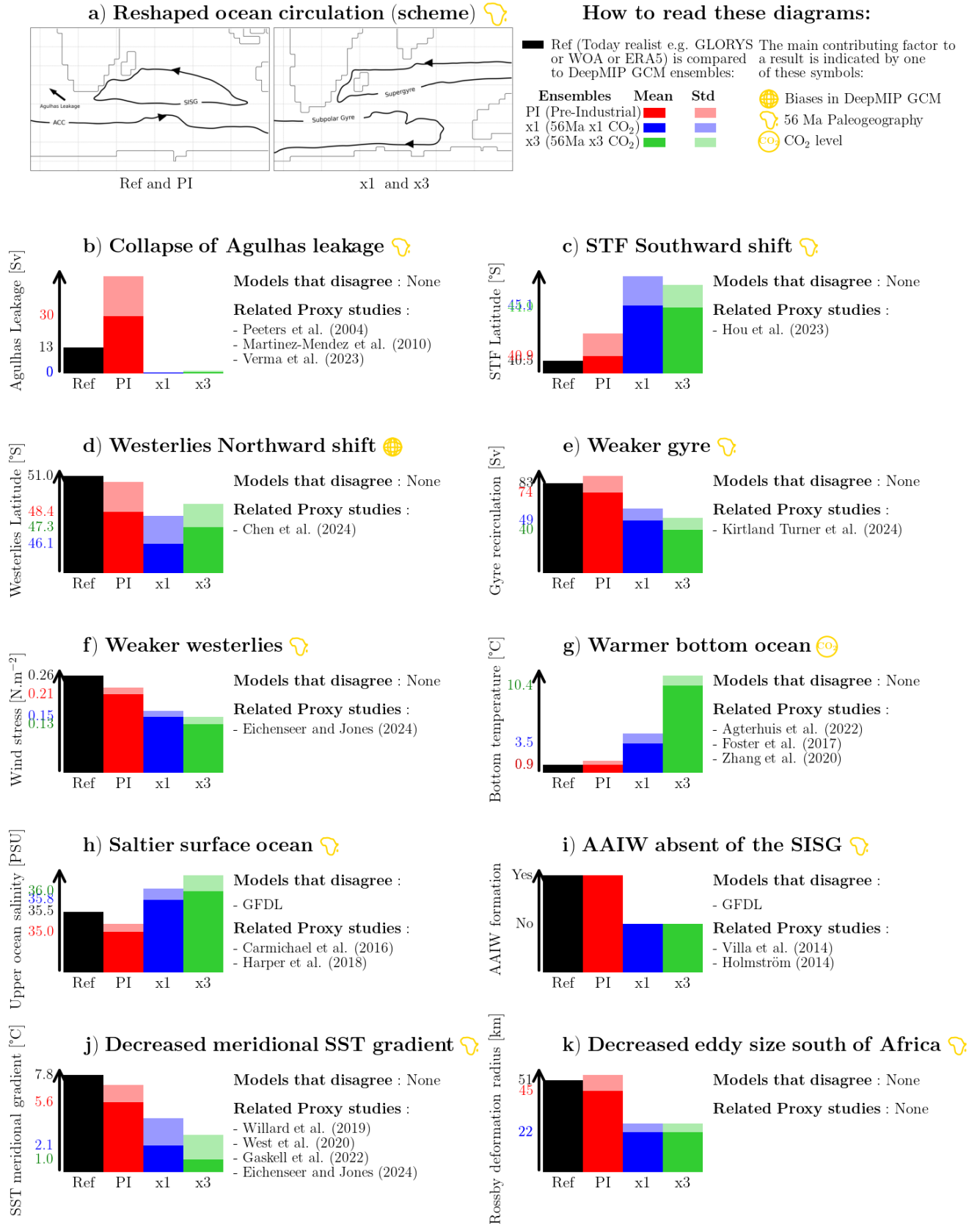


Figure 8: Summary of the main changes found in the Eocene DeepMIP simulations, the major contributing factor (highlighted in yellow), the models that disagree and related proxy studies. The values in the "Warmer bottom ocean" panel g) correspond to the average bottom water temperature in the characteristic gyre region (40°S–20°S, 50°E–100°E). For "Saltier surface ocean" panel h) the values represent the maximum salinity from the salinity profiles shown in Figure 6b. For "Decreased eddy size south of Africa" the values correspond to the Rossby deformation radius at 40°S/30°E for PI and 50°S/20°E for paleo x1 CO₂ and paleo x3 CO₂.

Barotropic streamfunction: 2 different methods

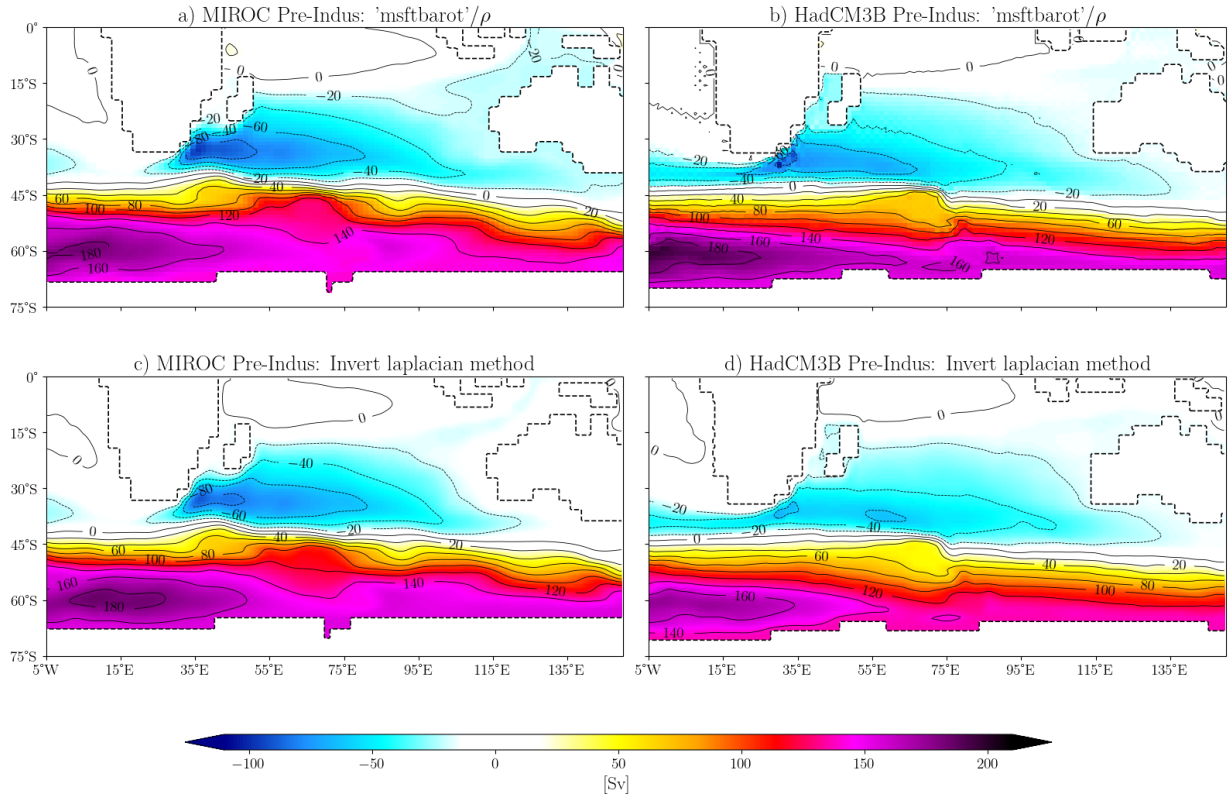


Figure A1: Barotropic streamfunction for the pre-industrial (PI) simulations of the MIROC and HadCM3B models, computed using two different methods: derived from the “msftbarot” in panels a) and b), and obtained by inverting a Laplacian in panels c) and d).

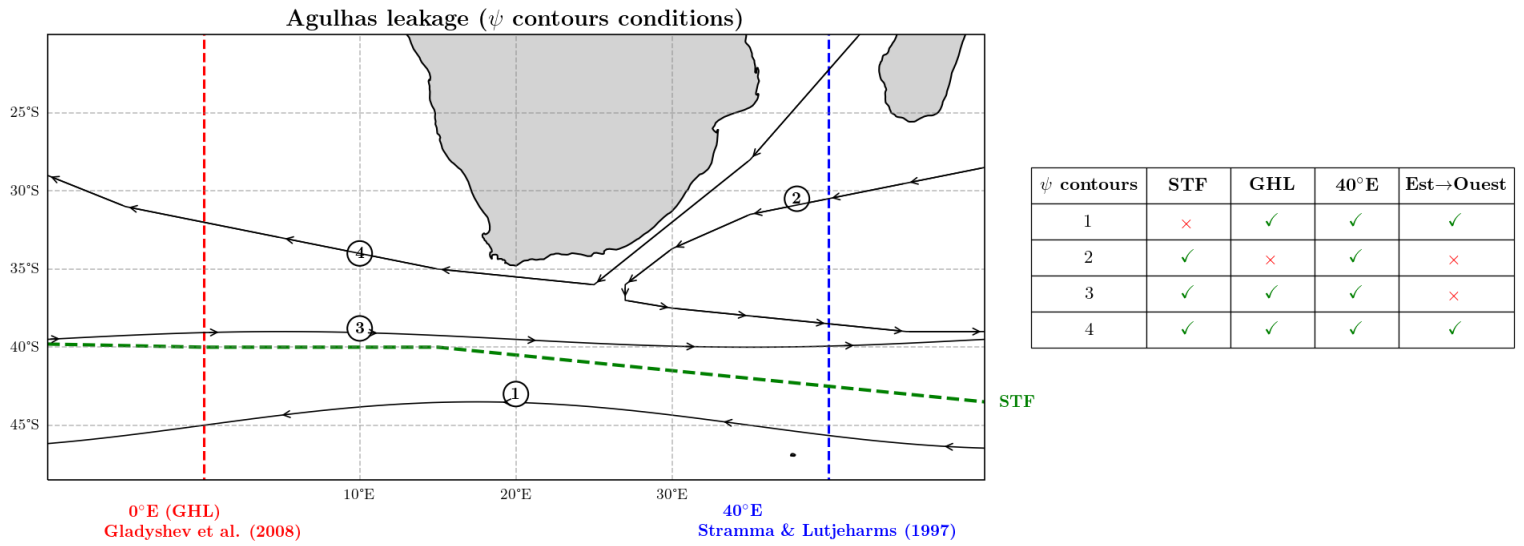


Figure B1: Schematic representation of the barotropic streamfunction (ψ) contours used to compute Agulhas leakage. Four example contours are shown, each representing different combinations of the conditions required to qualify as leakage pathways. These conditions include being northward of the Subtropical Front (STF), connecting to the Good Hope Line (GHL), connecting the 40°E meridional section, and exhibiting an east-to-west direction from the Indian Ocean to the Atlantic. Contours are labeled (1–4), with arrows indicating flow direction. The table on the right summarizes which conditions are fulfilled by each contour.