Energy Transfers between Multidecadal and Turbulent Variability

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ABSTRACT: One of the proposed mechanisms to explain the multidecadal variability observed in sea surface temperature of the North Atlantic Ocean consists of a large-scale low-frequency internal mode spontaneously developing because of the large-scale baroclinic instability of the time-mean circulation. Even though this mode has been extensively studied in terms of the buoyancy variance budget, its energetic properties remain poorly known. Here we perform the full mechanical energy budget including available potential energy (APE) and kinetic energy (KE) of this internal mode and decompose the budget into three frequency bands: mean, low frequency (LF) associated with the large-scale mode, and high frequency (HF) associated with mesoscale eddy turbulence. This decomposition allows us to diagnose the energy fluxes between the different reservoirs and to understand the sources and sinks. Because of the large scale of the mode, most of its energy is contained in the APE. In our configuration, the only source of LF APE is the transfer from mean APE to LF APE that is attributed to the large-scale baroclinic instability. In return the sinks of LF APE are the parameterized diffusion, the flux toward HF APE, and, to a much lesser extent, the flux toward LF KE. The presence of an additional wind stress component weakens multidecadal oscillations and modifies the energy fluxes between the different energy reservoirs. The KE transfer appears to only have a minor influence on the multidecadal mode relative to the other energy sources involving APE, in all experiments. These results highlight the utility of the full APE–KE budget.

KEYWORDS: Ocean dynamics; Energy budget/balance; Oceanic variability

1. Introduction

The multidecadal large-scale variability of the sea surface temperature (SST) is characterized in the North Atlantic by an anomaly intensified in the subpolar region and by a weaker anomaly of opposite sign south of the equator (Kushnir 1994; Deser et al. 2010; Zhang et al. 2019). This large-scale SST variability has been named Atlantic multidecadal variability (AMV; Kushnir 1994; Schlesinger and Ramankutty 1994; Kerr 2000; Sutton et al. 2018). Cool AMV phases occurred in the 1900s to 1920s and 1960s to 1990s and warm phases occurred in the 1930s to 1960s and after 1995. These cool and warm phases have been shown to be associated with several regional climate impacts such as the Sahel Indian summer monsoon rainfall, Atlantic hurricanes frequency, summer climate over western Europe and North America (Zhang et al. 2019), and wave climate in the Atlantic and Pacific Oceans (Reguero et al. 2019). Observations moreover show that more heat is released from the North Atlantic Ocean to the atmosphere during the positive phase of the AMV (Gulev et al. 2013). Thus, understanding what controls the dynamics of this variability and its potential predictability has essential societal and economic implications (Sutton et al. 2018).

Several mechanisms have been proposed to explain the origin of the AMV and remain actively debated [see, e.g., the recent discussions in Clement et al. (2015), Zhang et al. (2016), and Clement et al. (2016)]. Some studies suggest a direct role of the atmosphere either via stochastic heat flux (Hasselmann 1976; Frankignoul and Hasselmann 1977; Clement et al. 2015) or via aerosol emissions (Booth et al. 2012),

while other studies (e.g., Sévellec and Fedorov 2013; Arzel et al. 2018) suggest a role for oceanic processes linked with the internal variability of the Atlantic meridional overturning circulation (AMOC). In this work we focus on improving our knowledge of the physics of internal ocean modes, which are one of the possible explanations for the AMV. At low resolution, internal interdecadal variability arises in rectangular flatbottomed single hemispheric basin forced by prescribed surface heat fluxes (Greatbatch and Zhang 1995; Huck et al. 1999). This internal variability is due to a large-scale baroclinic instability that gives rise to SST anomalies and to geostrophically induced meridional overturning circulation variability (Colin de Verdière and Huck 1999; te Raa and Dijkstra 2002). This mode of variability and its mechanism were also demonstrated to exist in global realistic configuration of an ocean general circulation model (OGCM) (Sévellec and Fedorov 2013), in idealized coupled models (Buckley et al. 2012; Jamet et al. 2016), in climate models (Muir and Fedorov 2017), and in observations (Frankcombe et al. 2008). It is also shown to produce maximum SST variance in the region where the AMV signature is observed (Arzel et al. 2018).

The mode can be damped in some models [such as in the study of Sévellec and Fedorov (2013)] and self-sustained in others (e.g., Huck et al. 2015). The damped or self-sustained nature of the mode depends on different parameters such as the topography (Winton 1997), the wind shape and strength, or the vertical and horizontal diffusion (Huck et al. 2001; Arzel et al. 2018). In the case of a damped mode, atmospheric stochastic forcing is needed to excite the mode. Frankcombe et al. (2009) showed that the introduction of a North Atlantic Oscillation (NAO)-type stochastic forcing leads to an amplitude of sea surface temperature variability comparable to

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observations. Arzel et al. (2018) studied the bifurcation structure of the mode in a realistic configuration forced by prescribed surface fluxes and showed that the mode becomes damped for eddy induced diffusivities larger than 600 m s⁻¹.

In addition to available potential energy fluxes associated with the large-scale instability mechanism, ocean mesoscales eddies have been shown to be at the origin of a spatiotemporal inverse cascade of kinetic energy (Arbic et al. 2014; Martin et al. 2020). This latter mechanism has been proposed to be central to the existence of interannual-to-decadal fluctuations of sea level anomalies and surface kinetic energy in globalscale eddying simulations (Penduff et al. 2011; Arbic et al. 2014; Sérazin et al. 2015, 2018; Martin et al. 2020) and to influence the AMOC variability (Grégorio et al. 2015; Leroux et al. 2018; Jamet et al. 2019). However, realistic and global-scale eddy-resolving simulations of multidecadal variability are still beyond reach because of the long time integration required to bring the circulation into near equilibrium with the weak interior diffusive vertical fluxes. Therefore, most of the studies devoted to this problem are based on simple box-model geometries (e.g., Spall 2008; Huck et al. 2015; Hochet et al. 2020). How the oceanic mesoscale turbulence influences the multidecadal mode that spontaneously develops under prescribed surface fluxes has been explored by Huck et al. (2015). These authors show that, in the presence of mesoscale turbulence, the primary mechanism driving multidecadalscale temperature fluctuations remains the large-scale baroclinic instability mechanism. The presence of a surface restoring boundary condition in Spall (2008) prevents the internal ocean mode from developing so that the wind-driven gyre circulation and subsequent mesoscale instabilities play a major role. The coexistence of the mode described above and mesoscale eddies was shown by Huck et al. (2015) using idealized simulations at eddy-resolving resolution. In such eddying configurations and using a frequency-domain approach, Hochet et al. (2020) have highlighted a nonlinear transfer of temperature variance from low to high frequencies: mesoscale eddies are a sink of temperature variance for the low-frequency mode. Hence low frequencies do not arise as the result of the mesoscale eddy field, as in Spall (2008) for instance, but instead draw their energy source from the large-scale stratification. Sévellec et al. (2020) have also shown a similar behavior using mooring data in the Southern Ocean, but on shorter time scale. However the use of temperature variance instead of an energetic framework, as in Arbic et al. (2014) or Sérazin et al. (2015, 2018) makes the comparison with results from these studies difficult. Temperature variance is up to a factor equal to the local definition of available potential energy (APE) ε_{APE} in a quasigeostrophic (QG) framework:

$$\varepsilon_{\rm APE} \approx \frac{1}{2} \frac{g^2 {\rho'}^2}{\rho_0 N^2},\tag{1}$$

where N is the Brunt–Väisälä frequency, g is the acceleration of gravity parameter, ρ' is the density anomaly, and ρ_0 is the reference density. However, the internal mode described in the literature cited above occurs in regions where isopycnals

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outcropping prevents the use of Eq. (1). Thus, diagnostics and budget of temperature variance do not permit identifying the sources and sinks of energy. In this article we seek to obtain the full energy budget [i.e., including both kinetic energy (KE) and available potential energy] for the low-frequency mode described for instance in Hochet et al. (2020) and to quantify the energy transfers associated with the time-mean flow and mesoscale eddy field. In particular, we want to compare the intensity and direction of the conversion between low-frequency (LF) and high-frequency (HF) KE and between LF and HF APE. Because the only source of KE in the buoyancy forced experiment of Hochet et al. (2020) is the APE/KE conversion terms, we also investigate the effect of a wind stress forcing to add a direct source of KE.

The main difficulty in obtaining the energy budget in different frequency bands lies in the APE decomposition. Indeed, contrary to the kinetic energy, the time decomposition of the full APE formula is not straightforward. Scotti and White (2014) circumvented this problem by computing the fluctuating APE as the difference between the APE for the total circulation (i.e., time mean and fluctuations) and the APE for the time-mean circulation. This idea has been applied by Zemskova et al. (2015) to an eddy-permitting ocean state estimate [from phase II of the Estimating the Circulation and Climate of the Ocean (ECCO2)] to decompose the APE and KE budget into time-mean and fluctuating components. More recently, the same method has been used by Zemskova et al. (2021) to study the influence of several wind intensities over the Southern Ocean on the time-mean and fluctuating components of the APE and KE budget. They found that the APE budget is not significantly affected by the surface wind stress and is mainly controlled by the surface buoyancy forcing. However, we will show that applying this method to our configuration leads to a spurious imprint of the time-independent surface heat flux forcing on the LF and HF APE reservoirs, whereas we would instead expect the energy flux associated with the steady forcing to be entirely imparted to the background stratification. We will thus develop an alternative method that will be used to decompose the APE into mean, low-frequency, and high-frequency parts.

The article is organized as follows: in section 2, we decompose the mean and low and high frequencies and derive the budget for APE and KE. In section 3, we give a description of the model configuration used in this study. In section 4, we describe the variability in three numerical simulations with idealized North Atlantic configuration and under prescribed surface heat flux forcing with different wind forcing intensities. In section 5, we apply the energy budgets described in section 2 on the simulation outputs. In section 6, we conclude and discuss the main findings.

2. Theory

In this section we derive the APE and KE budgets for the mean, LF, and HF circulations. We use a linear equation of state for the density ρ that is only a function of temperature: $\rho = \rho_0(1 - \alpha\theta)$, where θ is the temperature, $\alpha = 2 \times 10^{-4} \text{ K}^{-1}$ is

$$\frac{\partial \rho}{\partial t} + \mathbf{v} \cdot \nabla \rho = D + F, \qquad (2)$$

where $\mathbf{v} = (u, v, w)$ is the 3D velocity with u, v, and w being the zonal, meridional, and vertical velocities; D and F represent the dissipation and surface forcing of density, respectively, the latter being constant with time and zero below the surface (i.e., no penetrative radiation). The time-independent forcing is used here to keep the problem simple, and we therefore do not account for slow variation of the forcing linked for instance with climate change. Note that the generalization of the theory presented below to a time-dependent F and to penetrative radiation is straightforward.

To obtain a separate budget for the low- and high-frequency parts of the APE and KE we decompose each field *M* into mean and low- and high-frequency parts:

$$M = \overline{M} + M^{\rm LF} + M^{\rm HF},\tag{3}$$

where M^{LF} and M^{HF} are the low- and high-frequency parts of M, respectively, and \overline{M} is the time mean. The time mean is computed using

$$\overline{M} = \frac{1}{T} \int_{T} M \, dt, \tag{4}$$

where *T* is the time length over which the integral is computed so that M^{LF} and M^{HF} satisfy $\overline{M^{\text{HF}}} = \overline{M^{\text{LF}}} = \overline{M^{\text{LF}}} M^{\text{HF}} = 0$. To decompose into HF and LF we use a low-pass Butterworth filter (cutoff frequency given in the following section). The low-pass filter is represented by a tilde so that M^{HF} satisfies $\widetilde{M^{\text{HF}}} = 0$ and $\widetilde{M} = \overline{M} + M^{\text{LF}}$.

Using this decomposition in frequency bands on Eq. (2) gives the following evolution equations for the mean, LF, and HF of ρ :

$$\frac{\partial \overline{\rho}}{\partial t} = -\overline{\mathbf{v}} \cdot \nabla \overline{\rho} - \overline{\mathbf{v}^{\text{LF}} \cdot \nabla \rho^{\text{LF}}} - \overline{\mathbf{v}^{\text{HF}} \cdot \nabla \rho^{\text{HF}}} + \overline{D} + F, \qquad (5)$$

$$\frac{\partial \rho^{\rm LF}}{\partial t} = -\overline{\mathbf{v}} \cdot \nabla \rho^{\rm LF} - \mathbf{v}^{\rm LF} \cdot \nabla (\overline{\rho} + \rho^{\rm LF}) + \overline{\mathbf{v}^{\rm HF} \cdot \nabla \rho^{\rm HF}} - \overline{\mathbf{v}^{\rm HF} \cdot \nabla \rho^{\rm HF}} + \overline{\mathbf{v}^{\rm LF} \cdot \nabla \rho^{\rm LF}} + D^{\rm LF}, \quad \text{and}$$
(6)

$$\frac{\partial \rho^{\rm HF}}{\partial t} = -\overline{\mathbf{v}} \cdot \nabla \rho^{\rm HF} - \mathbf{v}^{\rm LF} \cdot \nabla \rho^{\rm HF} - \mathbf{v}^{\rm HF} \cdot \nabla (\overline{\rho} + \rho^{\rm LF} + \rho^{\rm HF}) + \widetilde{\mathbf{v}^{\rm HF} \cdot \nabla \rho^{\rm HF}} + D^{\rm HF}.$$
(7)

APE is obtained as the difference between potential energy (PE) and background potential energy (BPE). We derive

expressions for PE, BPE, and APE in sections 2a, 2b, and 2c below.

a. Potential energy

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Multiplying Eq. (5) by zg gives an equation for the mean PE:

$$\frac{\partial g z \overline{\rho}}{\partial t} = -\nabla \cdot (g \overline{\mathbf{v}} z \overline{\rho}) + g \overline{w} \overline{\rho} + g \overline{w^{\text{HF}} \rho^{\text{HF}}} + g \overline{w^{\text{LF}} \rho^{\text{LF}}} -\nabla \cdot [g z (\overline{v^{\text{HF}} \rho^{\text{HF}}} + \overline{v^{\text{LF}} \rho^{\text{LF}}})] + g z \overline{D} + g z \overline{F}, \quad (8)$$

Integrating Eq. (8) on the volume V of the basin and time averaging results in the following equation:

$$\frac{d\mathbf{P}\mathbf{E}^{\mathrm{MEAN}}}{dt} = \underbrace{\int_{V} g\overline{w}\,\overline{\rho}\,dV}_{-C(\mathbf{P}\mathbf{E}^{\mathrm{MEAN}},\,\mathbf{K}\mathbf{E}^{\mathrm{MEAN}})} + \underbrace{\int_{V} g\overline{w^{\mathrm{LF}}\rho^{\mathrm{LF}}}\,dV}_{-C(\mathbf{P}\mathbf{E}^{\mathrm{MEAN}},\,\mathbf{P}\mathbf{E}^{\mathrm{LF}})} + \underbrace{\int_{V} g\overline{w^{\mathrm{HF}}\rho^{\mathrm{HF}}}\,dV}_{-C(\mathbf{P}\mathbf{E}^{\mathrm{MEAN}},\,\mathbf{P}\mathbf{E}^{\mathrm{HF}})} + \underbrace{\int_{V} gz\overline{D}\,dV}_{D_{\mathrm{PE}}}.$$
(9)

Note that the volume integral of the forcing term multiplied by z disappears because z = 0 at the surface and \overline{F} is zero below the surface. We note C(A, B) as the conversion term from A to B with C(A, B) = -C(B, A), if C(A, B) > 0 then C(A, B) acts to increase B. The $C(PE^{MEAN}, KE^{MEAN})$, $C(PE^{MEAN}, PE^{LF})$, and $C(PE^{MEAN}, PE^{HF})$ are respectively the conversion of mean PE to mean KE, of mean PE to LF PE, and of mean PE to HF PE. The term D_{PE} is interpreted as the rate of conversion of internal energy to potential energy (e.g., Hughes et al. 2009). There is no potential energy in anomalies because $\overline{gz\rho^{LF}} = \overline{gz\rho^{HF}} = 0$, and therefore $PE^{LF} = PE^{HF} = 0$. However, following Zemskova et al. (2015), we attribute the two terms $\int_V g\overline{w^{HF}\rho^{HF}} dV$ and $\int_V g\overline{w^{LF}\rho^{LF}} dV$ to $C(PE^{MEAN}, PE^{HF})$ and $C(PE^{MEAN}, PE^{LF})$ so that the formal budgets for PE^{LF} and PE^{HF} are

$$\frac{dPE^{LF}}{dt} = -C(PE^{LF}, KE^{LF}) + C(PE^{MEAN}, PE^{LF})$$
(10)

and

$$\frac{dPE^{HF}}{dt} = -C(PE^{HF}, KE^{HF}) + C(PE^{MEAN}, PE^{HF}), \quad (11)$$

from which we deduce that $C(PE^{HF}, KE^{HF}) = C(PE^{MEAN}, PE^{HF})$ and $C(PE^{LF}, KE^{LF}) = C(PE^{MEAN}, PE^{LF})$.

b. Background potential energy

To obtain an equation for the BPE, we first define the reference level $z_r(\rho, t)$, which is a function of time and density and corresponds to the depth that ρ would have in the Lorenz state of minimum potential energy (Lorenz 1955), a state where



FIG. 1. Schematic showing how the reference depth z_r associated with a given density ρ is calculated. (left) Meridional section of the physical space where volume V_1 (colored in red) above the isopycnal ρ = const is shown. (right) Adiabatic rearrangement where the previous isopycnal is horizontal with the same volume of water V_1 above it; $z_r(\rho)$ is the depth of this isopycnal after the rearrangement.

isopycnal surfaces would be horizontal. Following Saenz et al. (2015), a simple relationship between ρ and its reference level z_r can be derived using the result that an adiabatic rearrangement of the fluid parcels conserves the volume:

$$\int_{V(\rho,t)} dV = \int_{z_r}^0 A(z_r) \, dz_r,$$
 (12)

where $V(\rho, t)$ is the volume of water parcels with density ρ' lower than ρ at time t and A(z) is the area of the ocean at depth z. With a nonlinear equation of state for density that depends on temperature, salinity, and pressure, the procedure to obtain z_r is complex and is described for instance in Saenz et al. (2015). However, in this work we use two assumptions that greatly simplify the calculation of z_r . The first is the assumption of a linear equation of state depending only on temperature. This implies that z_r is a function of density ρ [which is itself a function of temperature: $\rho = \rho_0(1 - \alpha\theta)$] and time t: $z_r = Z_r(\rho, t)$. The second assumption is the flat-bottom basin with vertical boundaries so that the basin area is independent of depth; that is, A(z) = A. Using Eq. (12), z_r is then simply

$$z_r(\rho,t) = -\frac{V(\rho,t)}{A}.$$
(13)

Note that with a depth-dependent ocean area, the reference depth can easily be obtained by solving Eq. (12). A schematic illustrating how the reference depth $z_r(\rho, t)$ is obtained using volume $V(\rho)$ is shown in Fig. 1. This reference depth can be used to rewrite the density ρ as a function of z_r such that $\rho(X, t) = \rho_r[Z_r(X, t), t]$, with $Z_r(X, t) = z_r[\rho(X, t), t]$. On this schematic and in this article, we denote as "physical space" the usual space described by X = (x, y, z) and t and denote as "reference depth z_r and t.

Figure 2 shows the reference depth z_r as a function of time for different values of temperature from a simulation that will be presented in a following section. The time variation of the function $z_r(\rho, t)$ cannot be neglected here because it varies by more than 500 m for the largest densities due to the presence of the large-scale, low-frequency mode. In Zemskova et al. (2015) z_r variations with time are small because their study is not focused on the same time scale as ours: their time mean is computed over 20 years and their temporal variability is made of interannual, seasonal, and shorter time scales while our focus is on decadal to multidecadal time scales. Note that the larger variations of reference depth at larger densities can be attributed to the stronger (weaker) ρ_r gradient at shallower (deeper) reference depth. Time variation of the reference depths due to seasonal variation and to the presence of eddies was also reported in Zemskova et al. (2015).

The background potential energy is defined as follows:

BPE =
$$\int_{V} gZ_r(X,t)\rho(X,t) \, dV = \int_{-H}^{0} Agz'_r \rho_r(z'_r,t) \, dz'_r,$$
 (14)

where *H* is the basin depth (so that AH = V) and where the last equality is obtained from Eq. (12) and describes the calculation of the BPE in the reference space. We now want to obtain separate budgets for the mean, LF, and HF BPE. In Zemskova et al. (2015) the BPE budget is computed by first calculating the BPE of the mean flow that they define as



FIG. 2. Time evolution (yr) of the reference depth (m) of several isotherms (°C) in the experiment without wind forcing.



FIG. 3. Schematic view of the transfer between the different reservoirs of LF, HF, and mean APE and KE, along with BPE. The KE dissipation is in blue, KE forcing is in red, transfers between reservoirs are in orange, and BPE–APE transfers due to forcing and parameterized diffusion are in green and pink, respectively. Names of all transfers are indicated near the corresponding arrow and are summarized in Table 1.

$$BPE_Z^{MEAN} = \int_V g\overline{\rho} z_r(\overline{\rho}) \, dV, \qquad (15)$$

where z_r is the reference depth associated with \overline{p} . The BPE of anomalies is then defined as the difference between the total BPE and the mean BPE:

$$BPE_{Z}^{\text{anomalies}} = \int_{V} [g\overline{\rho z_{r}(\rho, t)} - g\overline{\rho} z_{r}(\overline{\rho})] \, dV.$$
(16)

Using this method in our simulation leads to part of the timeindependent heat flux forcing being attributed to $BPE_Z^{anomalies}$. Indeed, in this framework, the BPE forcing by heat fluxes is

$$\int_{V} g\overline{F}[\overline{z_r(\rho,t)} - z_r(\overline{\rho})] \, dV, \tag{17}$$

where \overline{F} is the time-independent net heat flux at the surface. Because of the nonlinearity of the z_r function (and its time dependence) we have that

$$\frac{z_r(\rho_1) + z_r(\rho_2)}{2} \neq z_r \left(\frac{\rho_1 + \rho_2}{2}\right),\tag{18}$$

for two different densities $\rho_1 \neq \rho_2$. Thus, the term in Eq. (17) is nonzero and the time-independent forcing acts on the anomalies. However, an analysis in terms of density variance

shows that the time-independent heat flux only acts on timemean density (Hochet et al. 2020). Density anomalies are found to be forced only by the term $\overline{\rho' \mathbf{u}'} \cdot \nabla \overline{\rho}$, which is usually interpreted as the signature of a large-scale baroclinic instability (see Colin de Verdière and Huck 1999). The fact that the term in Eq. (17) is nonzero is thus at odds with this interpretation.

To circumvent this problem, we develop below an alternative method to separate the BPE in frequency bands. The readers not interested in the details of the BPE decomposition may skip this section and refer to Fig. 3, which gives an intuitive view of the transfer between the different reservoirs.

As in the previous section, ρ is first decomposed into three frequency bands (mean, low-frequency, and high-frequency) so that

$$\rho = \overline{\rho} + \rho^{\rm LF} + \rho^{\rm HF},\tag{19}$$

and $\rho(X, t) = \rho_r(z_r, t)$ can trivially be written as

$$\rho(z_r,t) = -\frac{1}{A} \frac{\partial}{\partial z_r} \int_{V(z_r)} \rho(X,t) \, dV, \tag{20}$$

(22)

where $V(z_r)$ is the volume of water with $z'_r > z_r$. This formula simply states that the average of ρ on z_r surfaces is ρ by definition of $z_r(\rho, t)$. Using Eq. (19) in Eq. (20) gives

$$\rho = -\underbrace{\frac{1}{A} \frac{\partial}{\partial z_r} \int_{V(z_r)} \overline{\rho} \, dV}_{\rho_r^{\text{MEAN}}(z_r, t)} \underbrace{-\frac{1}{A} \frac{\partial}{\partial z_r} \int_{V(z_r)} \rho^{\text{LF}} \, dV}_{\rho_r^{\text{LF}}(z_r, t)} \underbrace{-\frac{1}{A} \frac{\partial}{\partial z_r} \int_{V(z_r)} \rho^{\text{HF}} \, dV}_{\rho_r^{\text{HF}}(z_r, t)},$$
(21)

 $BPE = \int_{V} \overline{\rho g z_{r}(\rho, t)} \, dV = \int_{V} \underbrace{\overline{\rho_{r}^{MEAN}(z_{r}, t) g z_{r}(\rho, t)} \, dV}_{BPE^{MEAN}} + \underbrace{\int_{V} \overline{\rho_{r}^{LF}(z_{r}, t) g z_{r}(\rho, t)} \, dV}_{BPE^{LF}} + \underbrace{\int_{V} \overline{\rho_{r}^{HF}(z_{r}, t) g z_{r}(\rho, t)} \, dV}_{BPE^{HF}},$

where BPE^{MEAN} , BPE^{LF} , and BPE^{HF} are respectively the BPE associated with the mean, LF, and HF densities. The time evolution of the BPE is then the sum of the time evolution of the mean, LF, and HF BPE:

$$\frac{dBPE}{dt} = \frac{dBPE^{MEAN}}{dt} + \frac{dBPE^{LF}}{dt} + \frac{dBPE^{HF}}{dt}.$$
 (23)

The evolution equation for the mean, LF, or HF BPE is

$$\frac{d\mathsf{BPE}^*}{dt} = \int_V g \frac{\partial \rho_r^*}{\partial t} z_r \, dV + \underbrace{\int_V g \rho_r^* \frac{\partial z_r}{\partial t} \, dV}_{=0}, \qquad (24)$$

where the symbol * represents MEAN, LF, or HF. The second term of the rhs is zero as shown in Winters et al. (1995) and later in Tailleux (2009) because ρ_r^* is constant on z_r surfaces by construction [see Eq. (21)].

The three evolution equations for the mean, LF, and HF density [i.e., Eqs. (5), (6), and (7)] are averaged on z_r surfaces using the following formula:

$$\frac{\partial \rho_r^*}{\partial t}(z_r, t) = \frac{1}{A} \frac{\partial}{\partial z_r} \left[\int_{V(z_r)} \frac{\partial \rho^*}{\partial t} \, dV \right],\tag{25}$$

where the symbol * represents the mean, LF, or HF density. Using Eq. (25) with Eqs. (5), (6), and (7), and inserting in Eq. (24) leads to the following relations for the three BPE budgets:

$$\frac{d\text{BPE}^{\text{MEAN}}}{dt} = -C(\text{BPE}^{\text{MEAN}}, \text{BPE}^{\text{LF}}) - C(\text{BPE}^{\text{MEAN}}, \text{BPE}^{\text{HF}})$$

$$+F_{\rm BPE^{MEAN}} + D_{\rm BPE^{MEAN}},\tag{26}$$

$$\frac{dBPE^{LF}}{dt} = -C(BPE^{LF}, BPE^{HF}) + C(BPE^{MEAN}, BPE^{LF}) + D_{BPE^{LF}},$$
(27)

and

where ρ_r^{MEAN} , ρ_r^{LF} , and ρ_r^{HF} are the average of ρ^{MEAN} , ρ^{LF} , and ρ^{HF} , respectively, on z_r surfaces. This ensures that ρ_r^{MEAN} , ρ_r^{LF} , and ρ_r^{HF} are functions of z_r (and time), and this property will be useful to obtain an evolution equation for the BPE as will become clear below. The BPE is then decomposed as follows:

$$\frac{dBPE^{HF}}{dBPE^{HF}} = C(BPE^{LF}, BPE^{HF}) + C(BPE^{MEAN}, BPE^{HF}) + D_{DDP}^{HF}$$

$$\frac{dBPE^{HF}}{dt} = C(BPE^{LF}, BPE^{HF}) + C(BPE^{MEAN}, BPE^{HF}) + D_{BPE^{HF}},$$
(28)

where the conversion terms are

$$C(BPE^{LF}, BPE^{HF}) = g \int_{V} \overline{z_r(\rho, t)(v^{HF} \cdot \nabla \rho^{HF} - v^{HF} \cdot \nabla \rho^{HF})} - \overline{z_r(\rho, t)v^{LF} \cdot \nabla \rho^{HF}} \, dV,$$
(29)

$$C(\text{BPE}^{\text{MEAN}}, \text{BPE}^{\text{LF}}) = g \int_{V} \overline{z_r(\rho, t)} \mathbf{v}^{\text{LF}} \cdot \nabla \rho^{\text{LF}} - \overline{z_r(\rho, t)} \overline{\mathbf{v}} \cdot \nabla \rho^{\text{LF}} dV,$$
(30)

and

$$C(\text{BPE}^{\text{MEAN}}, \text{BPE}^{\text{HF}}) = g \int_{V} \overline{z_r(\rho, t)} \mathbf{v}^{\text{HF}} \cdot \nabla \rho^{\text{HF}}} - \overline{z_r(\rho, t)} \overline{\mathbf{v}} \cdot \nabla \rho^{\text{HF}}} \, dV$$
(31)

Note that we have used the following relation to obtain the conversion term formulas:

$$\int_{V} \overline{z_r(\rho, t) \mathbf{v}} \cdot \nabla(\overline{\rho} + \rho^{\mathrm{LF}} + \rho^{\mathrm{HF}}) dV$$
$$= \int_{V} \overline{z_r(\rho, t) \mathbf{v}^{\mathrm{HF}}} \cdot \nabla(\overline{\rho} + \rho^{\mathrm{LF}} + \rho^{\mathrm{HF}}) dV$$
$$= \int_{V} \overline{z_r(\rho, t) \mathbf{v}^{\mathrm{LF}}} \cdot \nabla(\overline{\rho} + \rho^{\mathrm{LF}} + \rho^{\mathrm{HF}}) dV = 0.$$
(32)

The dissipation of BPE (Hughes et al. 2009; Zemskova et al. 2015) for each frequency band * is

$$D_{\rm BPE^*} = \int_V g_{\overline{z_r(\rho, t)}D^*} dV, \qquad (33)$$

and the forcing of the mean BPE is

$$F_{\rm BPE^{MEAN}} = \int_{V} g \overline{z_r(\rho, t)} \,\overline{F} \, dV. \tag{34}$$

The advantage of this approach over that of Zemskova et al. (2015) is that the time-independent heat flux forcing is entirely contained in BPE^{MEAN} and that we have explicit equations for the densities associated with BPE^{LF} and BPE^{HF}.

In the three Massachusetts Institute of Technology GCM (MITgcm) configurations described in the following section (section 3), the diffusive processes increase the mean BPE $(D_{\text{BPE}} > 0)$ and the surface heat flux forcing acts to decrease it ($F_{\text{RPE}^{\text{MEAN}}} < 0$). Because the APE varies in opposition to the

BPE the above mentioned forcing and dissipation have respectively an increasing and decreasing impact on the APE. The sign of the forcing and dissipation of BPE is consistent with results from previous studies using BPE (Hughes et al. 2009; Zemskova et al. 2015).

c. Available potential energy

The mean APE budget is obtained as the difference between the mean PE [Eq. (9)] and the mean BPE [Eq. (26)] budgets:

$$\frac{dAPE^{MEAN}}{dt} = \frac{dPE^{MEAN}}{dt} - \frac{dBPE^{MEAN}}{dt}$$

$$= \underbrace{-C(APE^{MEAN}, APE^{LF})}_{= -C(PE^{MEAN}, PE^{LF}) + C(BPE^{MEAN}, BPE^{HF}) = -C(PE^{MEAN}, PE^{HF}) + C(BPE^{MEAN}, BPE^{HF})}_{= -C(PE^{MEAN}, KE^{MEAN})} + \underbrace{-D_{APE^{MEAN}}}_{= -D_{PE^{MEAN}}} + \underbrace{F_{APE^{MEAN}}}_{= -F_{BPE^{MEAN}}}$$
(35)

The conversion, dissipation, and forcing terms of PE and BPE are derived in the two previous sections. Because $APE^{MEAN} = PE^{MEAN} - BPE^{MEAN}$, the evolution terms of BPE^{MEAN} appear in the APE^{MEAN} budget with a minus sign. The term $D_{APE^{MEAN}}$ can then either be seen as the dissipation of mean APE or as the conversion between mean APE and mean BPE due to the time-mean diffusive flux. Note that $D_{PE^{MEAN}}$ does

not explicitly appear in the BPE budget but it can be argued (see Hughes et al. 2009) that it contributes to the BPE budget and is thus added here as part of $D_{\rm APE^{MEAN}}$. Similarly, $F_{\rm APE^{MEAN}}$ can be seen as the forcing of mean APE or as the conversion of (mean) BPE to APE due to surface heat flux.

Because there is no potential energy in anomalies, APE in anomalies is only made of BPE. The LF APE budget is then

$$\frac{dAPE^{LF}}{dt} = \frac{dPE^{LF}}{dt} - \frac{dBPE^{LF}}{dt} = \underbrace{-C(APE^{LF}, KE^{LF})}_{= -C(PE^{LF}, KE^{LF})} + \underbrace{C(APE^{MEAN}, APE^{LF})}_{= C(PE^{MEAN}, PE^{LF}) - C(BPE^{MEAN}, BPE^{LF})} \underbrace{-C(APE^{LF}, APE^{HF})}_{= C(BPE^{LF}, BPE^{HF})} + \underbrace{D_{APE^{LF}}}_{= -D_{BPE^{LF}}}.$$
(36)

Similarly, the HF APE budget is

$$\frac{dAPE^{HF}}{dt} = \frac{dPE^{HF}}{dt} - \frac{dBPE^{HF}}{dt} = \underbrace{-C(APE^{HF}, KE^{HF})}_{= -C(PE^{HF}, KE^{HF})} + \underbrace{C(APE^{MEAN}, APE^{HF})}_{=C(PE^{MEAN}, PE^{HF}) - C(BPE^{MEAN}, BPE^{HF})} + \underbrace{C(APE^{LF}, APE^{HF})}_{= -C(BPE^{LF}, BPE^{HF})} + \underbrace{D_{APE^{HF}}}_{= -D_{BPE^{HF}}}.$$
(37)

d. Kinetic energy

In this section, the budgets for the total and low-frequency kinetic energy are derived. The horizontal momentum equations are

$$\frac{\partial u}{\partial t} + \mathbf{v} \cdot \nabla u - f \mathbf{v} = -\frac{1}{\rho_0} \frac{\partial p}{\partial x} + V_u + F_u \quad \text{and} \tag{38}$$

$$\frac{\partial v}{\partial t} + \mathbf{v} \cdot \nabla v + f u = -\frac{1}{\rho_0} \frac{\partial p}{\partial y} + V_v, \qquad (39)$$

where *p* is the pressure, V_u and V_v are the viscous term in respectively the zonal and meridional direction, F_u is the zonal, time-independent forcing (we assume no meridional forcing), and *f* is the Coriolis parameter. Time averaging Eqs. (38) and (39), multiplying by $\rho_0 \overline{u}$ and $\rho_0 \overline{v}$, and summing give an equation for the local kinetic energy of the time-mean flow:

$$\frac{\rho_0}{2}\overline{\mathbf{v}}\cdot\nabla(\overline{u}^2+\overline{v}^2) = -\rho_0\overline{u}\nabla\cdot\overline{(\mathbf{v}^{\mathrm{HF}}u^{\mathrm{HF}}+\mathbf{v}^{\mathrm{LF}}u^{\mathrm{LF}})} - \rho_0\overline{v}\nabla\cdot\overline{(\mathbf{v}^{\mathrm{HF}}v^{\mathrm{HF}}+\mathbf{v}^{\mathrm{LF}}v^{\mathrm{LF}})} - \nabla\overline{\mathbf{v}}\,\overline{p} - g\overline{\rho}\,\overline{w} + \rho_0\overline{u}\overline{V_u} + \rho_0\overline{v}\overline{V_v} + \rho_0\overline{u}\overline{F_u}. \tag{40}$$

The Coriolis term does not play any part in the kinetic energy budget because the Coriolis acceleration is normal to the velocity and there is no contribution from the vertical velocity because the hydrostatic approximation is used and implies no vertical acceleration (Gregory and Tailleux 2011). Integrating Eq. (40) over the entire volume gives the budget for KE^{MEAN} :

$$\frac{d\mathrm{KE}^{\mathrm{MEAN}}}{dt} = C(\mathrm{APE}^{\mathrm{MEAN}}, \mathrm{KE}^{\mathrm{MEAN}}) - C(\mathrm{KE}^{\mathrm{MEAN}}, \mathrm{KE}^{\mathrm{LF}}) - C(\mathrm{KE}^{\mathrm{MEAN}}, \mathrm{KE}^{\mathrm{HF}}) + D_{\mathrm{KE}^{\mathrm{MEAN}}} + F_{\mathrm{KE}^{\mathrm{MEAN}}}, \tag{41}$$

(

where

$$C(APE^{MEAN}, KE^{MEAN}) = -\int_{V} g\overline{\rho} \,\overline{w} \, dV, \qquad (42)$$

$$C(\mathrm{KE}^{\mathrm{MEAN}}, \mathrm{KE}^{\mathrm{LF}}) = \int_{V} \overline{\rho_0 \overline{u} \nabla(\mathbf{v}^{\mathrm{LF}} u^{\mathrm{LF}}) + \rho_0 \overline{v} \nabla(\mathbf{v}^{\mathrm{LF}} v^{\mathrm{LF}})} \, dV,$$
(43)

$$C(\mathrm{KE}^{\mathrm{MEAN}}, \mathrm{KE}^{\mathrm{HF}}) = \int_{V} \overline{\rho_0 \overline{u} \nabla(\mathbf{v}^{\mathrm{HF}} u^{\mathrm{HF}}) + \rho_0 \overline{v} \nabla(\mathbf{v}^{\mathrm{HF}} v^{\mathrm{HF}})} \, dV,$$
(44)

$$D_{\mathrm{KE}^{\mathrm{MEAN}}} = \int_{V} \rho_0 \overline{u} V_u + \rho_0 \overline{v} V_v dV, \text{ and} \qquad (45)$$

$$F_{\rm KE^{MEAN}} = \int_{V} \rho_0 \overline{u} F_u \, dV. \tag{46}$$

Proceeding similarly for the LF and HF, we obtain the following budgets for KE^{LF} and KE^{HF}:

$$\frac{d\mathbf{K}\mathbf{E}^{LF}}{dt} = C(\mathbf{APE}^{LF}, \mathbf{K}\mathbf{E}^{LF}) + C(\mathbf{K}\mathbf{E}^{MEAN}, \mathbf{K}\mathbf{E}^{LF}) - C(\mathbf{K}\mathbf{E}^{LF}, \mathbf{K}\mathbf{E}^{HF}) + D_{\mathbf{K}\mathbf{E}^{LF}},$$
(47)

$$\frac{d\mathbf{K}\mathbf{E}^{\rm HF}}{dt} = C(\mathbf{A}\mathbf{P}\mathbf{E}^{\rm HF}, \mathbf{K}\mathbf{E}^{\rm HF}) + C(\mathbf{K}\mathbf{E}^{\rm MEAN}, \mathbf{K}\mathbf{E}^{\rm HF}) + C(\mathbf{K}\mathbf{E}^{\rm LF}, \mathbf{K}\mathbf{E}^{\rm HF}) + D_{\mathbf{K}\mathbf{E}^{\rm HF}}$$
(48)

$$C(APE^{LF}, KE^{LF}) = -\int_{V} g \overline{\rho^{LF} w^{LF}} \, dV, \qquad (49)$$

$$C(APE^{HF}, KE^{HF}) = -\int_{V} g \overline{\rho^{HF} w^{HF}} \, dV,$$
(50)

$$C(\mathrm{KE}^{\mathrm{LF}}, \mathrm{KE}^{\mathrm{HF}}) = -\int_{V} \overline{\rho_{0} u^{\mathrm{HF}} \nabla(\mathbf{u}^{\mathrm{HF}} u^{\mathrm{LF}}) + \rho_{0} u^{\mathrm{HF}} \nabla(\mathbf{u}^{\mathrm{HF}} u^{\mathrm{LF}})} \, dV,$$
(51)

$$D_{\mathrm{KE}^{\mathrm{LF}}} = \int_{V} \rho_0 \overline{u^{\mathrm{LF}} V_u^{\mathrm{LF}} + \rho_0 v^{\mathrm{LF}} V_v^{\mathrm{LF}}} \, dV, \text{ and} \qquad (52)$$

$$D_{\rm KE^{\rm HF}} = \int_{V} \rho_0 \overline{u^{\rm HF} V_u^{\rm HF} + \rho_0 v^{\rm HF} V_v^{\rm HF}} \, dV.$$
(53)

e. Practical calculation of the APE/KE budget using model outputs

In Table 1 we describe how each term of the KE and APE budget mapped in Fig. 3 is computed using results from the previous subsections. Because of the very long time scales of the LF variability (\sim 50 yr) it would require too much storage to resolve the HF terms. However, we show below that the HF budgets can be obtained as the residual of the well-resolved LF budgets. As explained above, APE forcing and dissipation can also be seen as conversion between BPE and APE reservoirs. This is shown in Fig. 3 by the addition of a BPE reservoir that exchanges energy with the three APE reservoirs.

The total conversion from APE to KE can be obtained from the time-mean advection of the temperature, which is an output of the model:

$$C(APE, KE) = \int_{V} gz \nabla \overline{\mathbf{v}\rho} \, dV = \underbrace{\int_{V} g\nabla(z \overline{\mathbf{v}\rho}) \, dV}_{=0} - \int_{V} g \overline{w\rho} \, dV.$$
(54)

Then $C(APE^{HF}, KE^{HF})$ can be deduced from the knowledge of $C(APE^{LF}, KE^{HF})$ and $C(APE^{MEAN}, KE^{MEAN})$:

TABLE 1. List of all	l terms in the energe	y budget shown	in Fig. 3.
		,	

Term	Definition			
KE and APE dissipation				
$D_{\mathrm{KE}^{\mathrm{LF}}}$	$\rho_0 \int_V \overline{(u^{\mathrm{LF}} V_u^{\mathrm{LF}} + v^{\mathrm{LF}} V_v^{\mathrm{LF}})} dV$			
$D_{\mathrm{KE}^{\mathrm{MEAN}}}$	$\rho_0 \int_V (\overline{u} \overline{V_u} + \overline{v} \overline{V_v}) dV$			
$D_{\mathrm{KE}^{\mathrm{HF}}}$	$-F_{\mathrm{KE}^{\mathrm{MEAN}}} - C(\mathrm{APE},\mathrm{KE}) - D_{\mathrm{KE}^{\mathrm{LF}}} - D_{\mathrm{KE}^{\mathrm{MEAN}}}$			
$D_{ m APE^{LF}}$	$-\int_{V} g \overline{z_r(\rho,t)} D^{\rm LF} dV$			
$D_{ m APE}^{ m MEAN}$	$\int_{V} g \overline{z} \overline{D} dV - \int_{V} g \overline{z_r(\rho, t)} \overline{D} dV$			
$D_{ m APE^{HF}}$	$-F_{APE^{MEAN}} - D_{APE^{LF}} - D_{APE^{MEAN}} + C(APE, KE)$			
KE and APE forcing				
$F_{\mathrm{KE}^{\mathrm{MEAN}}}$	$ ho_0 \int_V \overline{u} \overline{F_u} dV$			
$F_{ m APE^{MEAN}}$	$-\int_V g\overline{z_r(ho)}\overline{F}dV$			
Conversion between reservoirs				
$C(KE^{LF}, KE^{HF})$	$C(APE^{LF}, KE^{LF}) + C(KE^{MEAN}, KE^{LF}) + D_{KE^{LF}}$			
$C(KE^{MEAN}, KE^{LF})$	$\rho_0 \int_{-\infty} [\overline{u} \nabla \cdot \overline{(\mathbf{v}^{\mathrm{LF}} u^{\mathrm{LF}})} + \overline{v} \nabla \cdot \overline{(\mathbf{v}^{\mathrm{LF}} v^{\mathrm{LF}})}] dV$			
$C(KE^{MEAN}, KE^{HF})$	J_V $F_{\text{KE}^{\text{MEAN}}} + C(\text{APE}^{\text{MEAN}}, \text{KE}^{\text{MEAN}}) + D_{\text{KE}^{\text{MEAN}}} - C(\text{KE}^{\text{MEAN}}, \text{KE}^{\text{LF}})$			
$C(APE^{LF}, APE^{HF})$	$-C(APE^{LF}, KE^{LF}) + C(APE^{MEAN}, APE^{LF}) + D_{APE^{LF}}$			
$C(APE^{MEAN}, APE^{LF})$	$-\int g\overline{\rho^{\mathrm{LF}}}w^{\mathrm{LF}}dV - g\int \overline{z_r(\rho,t)}v^{\mathrm{LF}}\cdot\nabla\rho^{\mathrm{LF}} - \overline{z_r(\rho,t)}\overline{v}\cdot\nabla\rho^{\mathrm{LF}}dV$			
$C(APE^{MEAN}, APE^{HF})$	$-C(APE^{MEAN}, APE^{LF}) - C(APE^{MEAN}, KE^{MEAN}) + D_{APE^{MEAN}} + F_{APE^{MEAN}}$			
$C(APE^{MEAN}, KE^{MEAN})$	$-\int_V g\overline{ ho}\overline{w}dV$			
$C(APE^{LF}, KE^{LF})$	$-\int_{V}g\overline{ ho^{\mathrm{LF}}w^{\mathrm{LF}}}dV$			
$C(APE^{HF}, KE^{HF})$	$C(APE, KE) - C(APE^{MEAN}, KE^{LF}) - C(APE^{MEAN}, KE^{MEAN})$			
C(APE, KE)	$\int_{V} g z \overline{\nabla \mathbf{v} \rho} dV$			

$$C(APE^{HF}, KE^{HF}) = C(APE, KE) - C(APE^{LF}, KE^{HF})$$

$$-C(APE^{MEAN}, KE^{MEAN}).$$
 (55)

The term $D_{\text{KE}^{\text{HF}}}$ is obtained from the total KE budget:

$$D_{\mathrm{KE}^{\mathrm{HF}}} = -C(\mathrm{APE}, \mathrm{KE}) - F_{\mathrm{KE}^{\mathrm{MEAN}}} - D_{\mathrm{KE}^{\mathrm{MEAN}}} - D_{\mathrm{KE}^{\mathrm{LF}}}.$$
(56)

The term $D_{APE^{HF}}$ is obtained using the total APE budget:

$$D_{\text{APE}^{\text{HF}}} = C(\text{APE}, \text{KE}) - F_{\text{APE}^{\text{MEAN}}} - D_{\text{APE}^{\text{MEAN}}} - D_{\text{APE}^{\text{LF}}}.$$
(57)

Similarly, $C(KE^{LF}, KE^{HF})$ is obtained from the KE^{LF} budget, $C(KE^{MEAN}, KE^{HF})$ from the KE^{MEAN} budget, $C(APE^{LF}, APE^{HF})$ from the APE^{LF} budget, and $C(APE^{MEAN}, APE^{HF})$ from the APE^{MEAN} budget.

3. Model and configuration

We use the MITgcm (Marshall et al. 1997) in a rectangular flat-bottom basin with a Cartesian geometry on a β plane centered at 40°N. The zonal and meridional extents are respectively $L_x = 5000$ km and $L_y = 4500$ km, and the southern boundary is located 2000 km north of the equator. An eddy-permitting horizontal resolution of 20 km is used in both directions. This resolution is sufficient in Huck et al. (2015) to capture the main characteristics of the effect of eddy turbulence on low-frequency variability. The depth is H = 4500 m, and there are 40 levels on the vertical with grid spacing increasing from 10 m at the surface to 400 m at the bottom.

The ocean is forced by constant heat flux at the surface, decreasing linearly with latitude from 50 W m⁻² at y = 0 km to -50 W m⁻² at y = 4500 km, similar to Huck et al. (2015). Static instability is removed by strong vertical mixing of the water



FIG. 4. Analytical zonal wind stress as a function of latitude for $\tau_0 = 0 \text{ N m}^{-2}$ (blue line), $\tau_0 = 0.05 \text{ N m}^{-2}$ (orange line), and $\tau_0 = 0.1 \text{ N m}^{-2}$ (green line).

column. We use biharmonic horizontal eddy diffusivity with a uniform value of 10^{11} m⁴ s⁻¹ and Leith implicit viscosity. The vertical viscosity is $\nu_{\nu} = 10^{-3}$ m² s⁻¹. In this single hemisphere configuration, the strength of the meridional overturning circulation (MOC) is a strong function of the vertical diffusivity K_{ν} , in

agreement with the $K_{\nu}^{1/2}$ geostrophic scaling (Huang and Chou 1994). Here, we choose to use $K_{\nu} = 2 \times 10^4$ m² s⁻¹ corresponding to a MOC strength close to 10 Sv. Because the primary objective of this study is to establish and understand the full energy budget of the low-frequency mode in the configuration used in previous published articles (Huck et al. 1999, 2001, 2015; Hochet et al. 2020) our main experiment does not have wind forcing. However, we also perform two additional experiments with increasing wind forcing intensity to study the effect of a direct KE source on the energy budget. The zonal wind stress used in the two wind forcing experiments varies with latitude according to the following formula:

$$\tau_x(y) = \tau_0 \left[\frac{1}{4} \cos\left(\frac{y\pi}{L_y}\right) - \cos\left(\frac{2\pi y}{L_y}\right) \right],\tag{58}$$

where τ_0 is the wind stress amplitude. The meridional wind stress is zero. We chose to use a nonsymmetric zonal wind stress as it seems important to achieve a generic dynamical behavior of the double-gyre circulation (Berloff and McWilliams 1999). The three experiments use $\tau_0 = 0 \text{ N m}^{-2}$ (no wind forcing), $\tau_0 = 0.05 \text{ N m}^{-2}$ (intermediate wind), and $\tau_0 = 0.1 \text{ N} \text{ m}^{-2}$ (climatological wind) (Fig. 4). All three experiments are



FIG. 5. Volume average of the temperature spectrum as a function of frequency calculated from 5-days (orange line) and 50-days (blue line) average outputs for (top left) $\tau_0 = 0 \text{ N m}^{-2}$, (top right) 0.05 N m⁻², and (bottom) 0.1 N m⁻². The black vertical lines on each panel show the peak magnitude for each experiment, and the red vertical lines show the LF/HF 3.5-yr separation.



FIG. 6. First complex EOF calculated on 1-yr-averaged 3D temperature outputs (°C) of the $\tau_0 = 0$ N m⁻² run, accounting for 60% of the variability: (top left) real part and (top right) imaginary part of the SST. CEOF (middle left) real part and (middle right) imaginary part along the meridional section (longitude = 800 km) shown by a red line on the top left and right panels. Black contours show isotherms of the time-mean temperature. (bottom) Real (red solid) and imaginary (red dotted) parts of the principal component of the first CEOF. The blue line shows the time evolution of the APE in exajoules.

initialized with a state of rest, the spinup time is then 500 years, and the model is run for another 400 years to produce outputs to compute the diagnostics presented below.

4. Time-mean circulation and variability

In the following section we describe the time-mean circulation as well as the low- and high-frequency variability obtained for the range of surface wind stress forcing amplitudes mentioned above. Hochet et al. (2020) used exactly the same model parameters and configuration as the present study with zero wind stress forcing. The turbulent transfer of temperature variance in their study was shown to act as a source of temperature variance for frequencies higher than 1 per 3.5 years and a sink for smaller frequencies. We thus define the limit between low and high frequencies as being 3.5 yr. Although it is possible that this limit is altered by the surface wind stress that we use in the two other experiments, we keep the same definition of 1/3.5 years in all experiments to be able to compare the three configurations. We thus associate LF with multidecadal, decadal, and part of the interannual variability and HF with part of the interannual and eddy turbulence induced variability.

a. No wind forcing $\tau_0 = 0 N m^{-2}$

Note that the simulation used in this section (i.e., without wind forcing) is the same as that described in Hochet et al. (2020), and the description of the LF variability is reproduced below. In the absence of wind stress forcing, LF variability spontaneously develops with a significant and narrow peak frequency of $1/53 \text{ yr}^{-1}$ (Fig. 5). A detailed description of the variability developing in very similar geometries can be found for instance in Huck et al. (1999) and Huck et al. (2015). Here we will only give a short description of its main characteristics. Following Hochet et al. (2020), we use complex empirical orthogonal functions (CEOFs) to describe the LF variability of the three-dimensional temperature field. The CEOFs are calculated using 50-day average outputs on a 400-yr-long simulation. The period of 400 years is chosen to obtain a statistical equilibrium of the solution.



FIG. 7. (left) Sea surface height standard deviation (m) calculated from 50-days average and time mean (black contours), (center) vertical integral of the time-mean LF density variance (kg⁻² m⁻⁵), and (right) time mean of the vertical integral of the LF APE (J m⁻²) for (top) $\tau_0 = 0$, (middle) 0.05, and (bottom) 0.1 N m⁻². Note that the color scale is different for each panel.

Similar to the widely used empirical orthogonal functions, CEOFs are the eigenvectors of the complex covariance matrix of a complex temperature anomaly that is calculated using the Hilbert transform of the detrended temperature anomaly (von Storch and Zwiers 2001). The leading CEOF contains 60% of the temperature variance (Fig. 6). The temperature anomaly associated with a CEOF can then be reconstructed using the following formula:

$$\theta_{\text{CEOF}}(x, y, z, t) = \text{PC}_{\text{re}}(t)\text{CEOF}_{\text{re}}(x, y, z)$$
$$+ \text{PC}_{\text{im}}(t)\text{CEOF}_{\text{im}}(x, y, z), \tag{59}$$

where the subscripts re and im stand for the real and imaginary parts, respectively, and PC is the principal component of the corresponding CEOF. The APE is shown along with the real and imaginary part of the PC (Fig. 6). The phase of the leading CEOF is chosen to match that of the APE time variation. The APE is very well correlated with the real part of the PC and shows that APE multidecadal variations are linked with the SST pattern shown in the upper panel of Fig. 6. The low-frequency variability takes the form of a large-scale temperature anomaly, located mainly in the northwestern half and in the upper 500 m of the basin with an SST anomaly larger than 3 K at some locations (Fig. 6). The successive positions of the positive and negative temperature anomalies as shown in Fig. 6—with a negative center located at latitudes around 2500–3000 km and longitude 2000 km (opposite of the imaginary part), followed by a negative center for latitudes between 3000 and 3500 km and longitude around 2000 km (real part), then a negative center around latitude 4000 km and longitude 1500 km (imaginary part), and then a negative center in the northwestern corner (opposite of the real part)—indicate northwestward propagation of the temperature anomalies.

The sea surface height (SSH) varies together with the temperature anomalies of the leading CEOF (left column of Fig. 7). The amplitude of SSH anomalies (15–20 cm) is maximum along the western boundary current and in its eastward extension in the northern half of the basin. These values compare well to altimetric observations (Stammer 1997). The time mean of the vertical integral of the LF APE and of the LF density variance are shown in Fig. 7. The largest values for the LF APE are located along the northern boundary. Equation (22) shows that large values of LF APE (= -LF BPE) are associated with deep reference depth and thus outcropping of dense waters. The



FIG. 8. As in Fig. 6, but for an intermediate double-gyre wind forcing of $\tau_0 = 0.05$ N m⁻². The leading CEOF accounts for 28% of the temperature variance. The red line on the two top panels shows the same meridional section as in Fig. 6 used to plot the two panels in the middle.

location of LF APE contrasts with the location of the largest values of LF density variance, which are located in the northern part of the basin interior (Fig. 7). The differences between these two quantities further demonstrate that the APE cannot be approximated by the density variance in these configurations in contrast with QG theory.

b. Intermediate wind forcing $\tau_0 = 0.05 \ N \ m^{-2}$

When wind stress forcing is present the temperature variability is significantly reduced relative to the previous case and consists of a broad band of low-frequency signals with a peak frequency of ~1/22 yr⁻¹ (Fig. 5). Huck et al. (2001) explained this effect of the wind forcing on the low frequency variability by the damping effect of the Ekman pumping on the large-scale anomalies. To explain this effect, Huck et al. (2001) assumed that the following formula describes the effect of wind stress on the temperature anomalies [see their Eq. (14)]:

$$\frac{\partial \theta'}{\partial t} = -W_E \frac{\partial \theta'}{\partial z},\tag{60}$$

where θ' is the temperature anomaly and W_E is the Ekman pumping. Then if θ' is further assumed to have an exponential

profile with depth, $\theta' \propto \exp(-W_E kt)$ with k on the order of 500 m⁻¹. The temperature anomaly decreases where W_E is positive, which is in the northern half of the basin in our configuration.

The surface signature of the leading CEOF of temperature variability, explaining 28% of the spatially integrated variance, shows that the variability now occurs predominantly along the eastern boundary and along a narrow latitudinal band extending across the width of the basin just south of the intergyre boundary (Fig. 8). The large-scale anomaly emanates from the eastern boundary and propagates to the west along the mean temperature contours. SSH variability is no longer collocated with SST variations, as was the case with zero wind stress forcing, but instead mostly occurs along a region centered about the intergyre (at y = 2000 km) along the western boundary (Fig. 7, second row). The LF density variance is now almost entirely located on the eastern boundary as also shown by the CEOF (Fig. 8). The LF APE is in the northeastern corner of the basin in the region where dense waters outcrop.

c. Climatological wind forcing $\tau_0 = 0.1 N m^{-2}$

Increasing the amplitude of the zonal surface wind stress forcing up to realistic values has the effect of further



FIG. 9. As in Fig. 6, but for a climatological double-gyre wind forcing of $\tau_0 = 0.1$ N m⁻². The leading CEOF accounts for 7% of the temperature variance.

decreasing the temperature variance on interdecadal time scales and increasing the temperature variance on monthlyto-interannual time scales relative to the case with $\tau_0 = 0.05$ N m⁻². Indeed, the volume averaged LF (HF) temperature spectrum (defined as frequencies lower than 2 per 3.5 years) is weaker for $\tau_0 = 0.05$ N m⁻² than for $\tau_0 = 0.01$ N m⁻² (Fig. 5).

The leading CEOF of temperature variability now represents only 7% of the spatially integrated variance and is mostly apparent south of the intergyre boundary (Fig. 9). This pattern of variability differs from the previous case with $\tau_0 =$ 0.05 N m⁻² (Fig. 8) for which SST variability was also present along the eastern boundary. SSH variability however shares the same pattern and amplitude as that obtained for $\tau_0 = 0.05$ N m⁻² with enhanced variability along the western boundary current. The LF APE is now located in the northwestern part of the basin with a much smaller amplitude than in the two previous cases.

5. Energy budget

We now describe the mean, LF, and HF KE, APE, and BPE budgets for the three experiments described above (summarized in Figs. 10, 12, and 13 following the schematic

given in Fig. 3). Table 2 gathers the transfer values obtained for the three experiments.

a. No wind forcing $\tau_0 = 0 N m^{-2}$

In the absence of surface wind stress there is no external source for the KE reservoir and the main energy pathways are located within the APE part of the budget (Fig. 10). Of the 119 GW of conversion between BPE and APE due to surface heat flux, 32 GW are converted into KE (mainly at HF) where it is dissipated by viscous forces. The remaining 87 GW are mainly transferred from HF APE to BPE because of dissipation (53 GW from APE HF to BPE, 16 GW from LF APE to BPE, and 18 GW from mean APE to BPE).

Despite the differences between BPE and density variance shown in Fig. 7, Hochet et al. (2020) found a similar pathway for the temperature variance. The surface heat flux is the only source of mean temperature variance and LF temperature variance is forced through a transfer of temperature variance from the mean flow; this transfer is interpreted as the result of baroclinic instability of the mean flow. The conversion from mean APE to LF APE is 49 GW, which is the largest conversion term in this experiment. The main sink of LF APE is the conversion to HF APE (29 GW), whereas the dissipation removes 16 GW. Because the direction of the LF/HF APE



FIG. 10. Schematic showing the exchange of energy between the different reservoirs (shown by black boxes) for the $\tau_0 = 0$ N m⁻² run. The conversion from BPE to APE forcing achieved by surface heat fluxes is shown by green arrows, and the conversion of APE to BPE because of diffusive fluxes is shown in pink. Dissipation of KE by viscous forces is shown by blue arrows, and the conversion between the different reservoirs is shown by orange arrows. The direction of the conversion follows the arrow direction. Conversions are expressed in gigawatts (=10⁹ W), the KE is in petajoules (=10¹⁵ J), and the APE is in exajoules (=10¹⁸ J).

transfer is from LF to HF and because LF is associated with large scales and HF with mesoscale eddies, we deduce that mesoscale eddy turbulence is a sink of energy for the low-frequency variability, which is one of the main results of this study. It confirms previous findings of Arbic et al. (2014), who demonstrated the existence of a direct temporal APE cascade along with the inverse temporal KE cascade under QG approximation. Conversion between kinetic energy reservoirs is small relative to conversion between APE reservoirs. There is however a substantial energy transfer (20 GW) between HF APE and HF KE. This input of HF KE is balanced by the sink linked with viscous terms. The only source of mean KE is the conversion of mean APE to mean KE (8 GW) and the conversion between LF and HF KE is negligible. The ratio of LF KE to LF APE is of 0.3%, showing that the low-frequency variability is predominantly found in APE in this simulation.

The two left columns of Fig. 11 show the spatial pattern of $C(APE^{MEAN}, APE^{HF})$ and $C(APE^{MEAN}, APE^{LF})$ integrated vertically over the water column. Strong positive values of these two terms are generally located close to the northern boundary, where the convection is the strongest. Positive values of $C(APE^{MEAN}, APE^{HF})$ are located in the eastern part of the northern boundary and follows closely the

values of mean APE forcing (last column of Fig. 11). Negative values are located in the southern part of the basin interior. The mean APE forcing term is large in regions where the reference level is the deepest (i.e., where dense waters outcrop at the surface). Note that $C(APE^{MEAN}, APE^{LF})$ (Fig. 11, middle-left column) has its largest values close to the northern and eastern boundary. The vertical integral of the term $-\overline{\mathbf{u}^{\mathrm{LF}}\rho^{\mathrm{LF}}}\cdot\nabla\overline{\rho}$, which represents the transfer from the mean to the LF density variance (see, e.g., Colin de Verdière and Huck 1999), is shown in the third column of Fig. 11 and is very different from $C(APE^{MEAN}, APE^{LF})$. This could be expected from the difference between the variance of the LF density and the LF APE already shown in Fig. 7: the transfer from mean APE to LF APE occurs in the convection region where dense waters outcrop whereas the LF density variance transfer occurs in the basin interior. We do not show $C(APE^{LF}, APE^{HF})$, but it follows closely the variation of $C(APE^{MEAN}, APE^{LF}).$

b. Intermediate wind forcing $\tau_0 = 0.05 N m^{-2}$

With an intermediate zonal wind stress at the surface, the LF variability becomes weaker and shifts to midlatitudes as explained in section 4b. The energy budget (Fig. 12) shows a



FIG. 11. Vertical integral of the conversion between (left) mean APE and HF APE [$C(APE^{MEAN}, APE^{HF})$] and (left center) mean APE and LF APE [$C(APE^{MEAN}, APE^{LF})$] (GW), (right center) the vertical integral of the density variance transfer from mean to LF (kg² m⁻⁶ s⁻¹; note that the range of the color bar is different in the top panel), and (right) the mean APE forcing (or transfer from BPE to mean APE due to heat fluxes) (GW) for the three wind intensities (top) $\tau_0 = 0$ N m⁻², (middle) $\tau_0 = 0.05$ N m⁻², and (bottom) $\tau_0 = 0.1$ N m⁻². Note that $C(APE^{LF}, APE^{HF})$ is not shown but looks very similar to $C(APE^{MEAN}, APE^{LF})$.

decrease of the energy fluxes from mean APE and LF APE. The transfer of BPE to mean APE is the same as that obtained for the no-wind experiment (119 GW). Part of this mean APE energy input is transferred back to BPE by parameterized diffusive flux (30 GW) whereas most of it is converted into HF APE (61 GW). The conversion of LF APE to HF APE is approximately half the value obtained for the no-wind case (14 vs 29 GW). The total transfer (i.e., KE + APE) is still directed from LF toward HF. There is now a small direct forcing of mean KE (7 GW) that adds to the conversion from mean APE to mean KE (5 GW) to create a source of 12 GW of mean KE. A total of 8 GW is directly dissipated by viscous forces, and the remainder is mainly transferred to HF KE. The conversion between LF KE and HF KE is negligible.

The ratio of LF KE and LF APE has increased when compared with the no-wind simulation (the ratio is now 1%); however, LF KE remains negligible relative to LF APE. We explain the predominance of APE over KE in the LF by the larger scales found at these frequencies. It is indeed known from QG theory that the ratio KE over APE decreases with larger scales (Vallis 2017). Using the ECCO2 ocean state estimate, Zemskova et al. (2015) found that there is approximately 10 times more APE than KE in anomalies, but they do not discriminate between frequency bands, which makes the comparison with our results difficult. With wind forcing, the $C(APE^{MEAN}, APE^{HF})$ conversion increases by 17 GW (Table 2); its positive values follow the northern boundary of the basin (Fig. 11) and the mean APE forcing (last column of Fig. 11). The largest values of $C(APE^{MEAN}, APE^{LF})$ are now almost entirely located in the northeastern corner where the SST anomalies seem to originate.

c. Climatological wind forcing $\tau_0 = 0.1 N m^{-2}$

In this experiment the zonal wind forcing is twice as strong as in the previous experiment, close to the climatological amplitude. The wind stress adds 34 GW to the mean KE reservoir and 15 GW is directly dissipated by viscous forces (Fig. 13). The conversion from mean APE to mean KE is now negative (i.e., from mean KE to mean APE), which is in line with what is calculated in OGCM (e.g., Toggweiler and Samuels 1998; Gnanadesikan et al. 2005; Gregory and Tailleux 2011). The remaining 8 GW are all converted to HF KE. The conversion of BPE to mean APE due to heat flux has slightly increased relative to the two other experiments (128 GW). The BPE-to-mean-APE conversion formula $-g \int_{U} z_r \overline{(\rho, t)} \overline{F} dV$ shows that the value of this conversion mainly depends on the position of the deepest reference depths and thus on the circulation in the northern half of the basin. An explanation for the (small) increase in BPE-tomean-APE conversion could thus be that only the



FIG. 12. As in Fig. 10, but with an intermediate double-gyre wind stress surface forcing of $\tau_0 = 0.05$ N m⁻². The KE forcing made by the wind stress is shown with a red arrow.

climatological wind substantially modifies the circulation in this region. The conversion from mean APE to HF APE is now much larger than the conversion from mean APE to LF APE (84 vs 13 GW), and 42 GW is directly dissipated. The energy in the LF KE reservoir has increased relative to the two other simulation and now represents 10% of LF APE.

With this realistic amplitude of wind forcing, positive values of the conversion $C(APE^{MEAN}, APE^{HF})$ continue to extend along the northern boundary (Fig. 11 bottom line) and is particularly intense in the northwestern corner. The $C(APE^{MEAN}, APE^{LF})$ intensity is weaker than before and almost entirely located in the northwestern corner.

Sohail et al. (2018) and Zemskova et al. (2021) report that the input of KE obtained by increasing the wind stress over the Southern Ocean results mainly in an increase of the KE dissipation term, rather than an increase in APE dissipation. These results are in line with what is found in our experiments: the total dissipation of KE $D_{\text{KE}^{\text{MEAN}}} + D_{\text{KE}^{\text{LF}}} + D_{\text{KE}^{\text{HF}}}$ has increased by 65% between the no-wind and climatological-wind experiments and by 24% for the total dissipation of APE $D_{\text{APE}^{\text{MEAN}}} + D_{\text{APE}^{\text{LF}}} + D_{\text{APE}^{\text{HF}}}$: most of the additional mean KE forcing is directly dissipated by KE total dissipation.

6. Conclusions

In this article we have derived the mechanical energy budget for the large-scale, internally generated, low-frequency ocean mode that was studied extensively in previous works (Colin de Verdière and Huck 1999; Huck et al. 1999; Huck and Vallis 2001; Huck et al. 2001, 2015; Arzel et al. 2018; Hochet et al. 2020). The mechanical energy budget is decomposed into mean, low-frequency, and high-frequency parts to study the effect of the large-scale baroclinic instability of the mean circulation and the effect of the eddy field on the LF mode. One of the main achievements of this work is the new way of decomposing the background potential energy into frequency bands that allows us to correctly attribute the source and sink terms associated with each reservoir.

The energy budget of the no-wind experiment shows that the energy of this multidecadal mode is mostly contained in the LF APE rather than in the LF KE because of its large scale. Using ECCO2, Zemskova et al. (2015) also found that the energy is mostly contained in the fluctuating APE rather than in the fluctuating KE, but with a smaller ratio of approximately 10 (as compared with an LF-APE-over-LF-KE ratio of approximately 300 in our no-wind experiment). This difference is probably due to our use of a single hemispheric basin and thus to the omission of the strong wind forcing over the Southern Ocean. The source term for the LF APE is the conversion from the mean APE to the LF APE. In agreement with Hochet et al. (2020) where the budget was made in terms of temperature variance, the sink term of the mode is attributed to the parameterized diffusion (\sim 36%) and to the transfer of APE to higher frequencies linked with mesoscale eddy turbulence (~64%). In the



FIG. 13. As in Fig. 10, but with a double-gyre wind stress surface forcing of $\tau_0 = 0.1 \text{ N m}^{-2}$. The KE forcing made by the wind stress is shown with a red arrow.

experiment with no wind forcing, the only source of kinetic energy is the conversion from APE at all frequencies. The transfer of KE between high and low frequencies, which is shown to be an important source of low-frequency variability in other experiments (Arbic et al. 2014; Sérazin et al. 2018), appears negligible in our configuration relative to all of the other energy fluxes. Recognizing that this might be due to the absence of any direct source of kinetic energy we performed two other experiments with a time-independent zonal wind forcing at the surface that drives the classical wind-driven double gyres.

With the addition of a wind forcing at the surface, a source term for the mean kinetic energy appears. The structure of the LF mode is modified with intermediate wind strength and almost disappears with climatological wind. With increasing wind forcing, the LF APE and all its associated conversion terms decrease. Indeed there is a decrease in the values of the conversion terms from mean APE to LF APE and of LF APE to BPE (linked with diffusive flux), of LF APE to HF APE, and of LF APE to LF KE for the sink terms. Meanwhile, the energy in the HF APE increases as well as the conversion from mean APE to HF APE. This larger conversion is balanced mostly by a larger conversion of HF APE to BPE due to diffusive flux. The energy contained in the mean APE increases as well as the conversion from BPE to mean APE due to heat flux. This increase is balanced by a larger conversion to BPE due to diffusive flux and by a larger conversion

to HF APE. The conversion of mean APE to mean KE becomes negative, which is in agreement with what is usually calculated in more realistic models such as in Zemskova et al. (2015). The conversion between LF KE and HF KE remains negligible or very small relative to other conversions; none-theless, it is directed from HF to LF for the climatological-wind experiment, which is in agreement with the temporal inverse KE cascade found in Arbic et al. (2014).

For all wind stress intensities studied here, the energy and conversion terms remain mainly contained in the APE. There is however a nonnegligible transfer of APE to KE at HF where it is dissipated by viscous forces. The fact that most of the transfers of energy occur between the different APE and BPE reservoirs rather than between PE or KE reservoirs outline the importance of the APE budget to study large-scale and low-frequency variability. In contrast, Arbic et al. (2014) and Sérazin et al. (2015, 2018) found a predominant role for the temporal inverse KE cascade, but we attribute this difference for the most part to our focus on multidecadal variability as compared with their focus on shorter interannual variability. The transfer from LF APE to HF APE demonstrates the damping role of the mesoscale eddy turbulence for the largescale variability, even for realistic wind intensities. This transfer of APE from LF to HF is very similar to the QG APE direct temporal cascade of APE that has been observed (along with the inverse temporal cascade of KE) in the idealized simulation of Arbic et al. (2014). Temperature variance

TABLE 2. Transfer values for all terms in the energy budget for the no-wind ($\tau_0 = 0 \text{ N m}^{-2}$), intermediate-wind $\tau_0 = (0.05 \text{ N m}^{-2})$),
and climatological-wind ($\tau_0 = 0.1 \text{ N m}^{-2}$) experiments.	

	$\tau_0 = 0 \text{ N m}^{-2}$	$\tau_0 = 0.05 \text{ N m}^{-2}$	$\tau_0 = 0.1 \text{ N m}^{-2}$
KE and APE dissipation			
$D_{\rm KE^{\rm LF}}$	-5 GW	-3 GW	-2 GW
$D_{\rm KE^{MEAN}}$	-5 GW	-8 GW	-15 GW
$D_{\rm KE^{\rm HF}}$	-22 GW	-22 GW	-36 GW
$D_{APE^{LF}}$	-16 GW	-7 GW	-5 GW
$D_{APE^{MEAN}}$	-18 GW	-30 GW	-42 GW
$D_{ m APE}^{ m HF}$	-53 GW	-56 GW	-62 GW
KE and APE forcing			
F _{KE^{MEAN}}	0 GW	7 GW	34 GW
$F_{APE^{MEAN}}$	119 GW	119 GW	128 GW
Conversion between reservoirs			
$C(KE^{LF}, KE^{HF})$	0 GW	0 GW	-1 GW
$C(KE^{MEAN}, KE^{LF})$	1 GW	1 GW	0 GW
$C(KE^{MEAN}, KE^{HF})$	2 GW	3 GW	8 GW
$C(APE^{LF}, APE^{HF})$	29 GW	14 GW	7 GW
$C(APE^{MEAN}, APE^{LF})$	49 GW	23 GW	13 GW
$C(APE^{MEAN}, APE^{HF})$	44 GW	61 GW	84 GW
$C(APE^{MEAN}, KE^{MEAN})$	8 GW	5 GW	-11 GW
$C(APE^{LF}, KE^{LF})$	4 GW	2 GW	1 GW
$C(APE^{HF}, KE^{HF})$	20 GW	19 GW	29 GW
C(APE, KE)	32 GW	26 GW	19 GW

budget gives similar pathway for the sources and sinks of the low frequency mode (Hochet et al. 2020). However, we have shown that the locations of the APE and of its associated transfers differ significantly from that of the density variance. Large values of APE are indeed linked with the outcropping of dense waters at the surface due to convection that occurs along the northern boundary in our configuration, whereas there is no significant link between large values of density variance and convective regions. The APE budget is therefore more accurate in identifying regions where energy conversions are the most important.

This study is limited by several approximations, the first being the representation of the North Atlantic Ocean with a simplified geometry and a flat bottom. In particular, the wind forcing over the Southern Ocean is thought to play an important role in setting the mean stratification, in particular the stratification in the North Atlantic (e.g., Nikurashin and Vallis 2011, 2012), and might therefore influence the dynamic of the low-frequency mode. Moreover, using the energy budget of a global ocean-sea ice model Hogg et al. (2017) have shown that wind change over the Southern Ocean leads to change of APE and stratification in the North Atlantic. However, it has been shown in Arzel et al. (2007) that the addition of a reentrant channel representing the Antarctic Circumpolar Current acts to reduce the low-frequency variability in the Southern Hemisphere but does not suppress the internal variability in the Northern Hemisphere. Moreover, the physical mechanism giving rise to the internal mode studied in this single hemispheric configuration is also found to give rise to low-frequency variability in a realistic ocean-only configuration (e.g., Sévellec and Fedorov 2013; Arzel et al. 2018; Arzel and Huck 2020) and an ocean-atmosphere coupled configuration (e.g.,

Ortega et al. 2015; Gastineau et al. 2018). The omission of salinity and the use of a linear equation of state for density certainly has an influence on the APE budget. Indeed, it is known that nonlinearities of the equation of state are in general not negligible (e.g., Klocker and McDougall 2010; Nycander et al. 2015). Nonetheless, it has been shown in a realistic setup (Sévellec and Fedorov 2013) that the mode is largely controlled by temperature variation in the upper ocean. Our study, as well as previous studies of the internal mode (e.g., Sévellec and Fedorov 2013; Huck et al. 2015; Arzel et al. 2018), assumes that external forcing is either constant or made of natural variability such as the North Atlantic Oscillation (Frankcombe et al. 2009; Arzel and Huck 2020). This assumption is helpful to understand the physics and the mechanisms of the mode. However, with its multidecadal variability, the internal mode might be affected by anthropogenic forcing, which acts on the same time scales and modifies the characteristics of the ocean stratification (Levitus et al. 2012).

The eddy-permitting resolution of 20 km used here is not sufficient to entirely resolve the eddy field; however, similar experiments at 10 km with no wind forcing were conducted in Huck et al. (2015) and no qualitative differences were found. Last we set the limit between LF and HF to be 3.5 years based on the results from Hochet et al. (2020) that showed that in the same configuration without wind forcing, nonlinear transfers of temperature variance are a source (sink) term for periods longer (shorter) than 3.5 years. To be able to compare the three configurations studied in this article we kept this limit fixed; however, with the addition of a wind forcing at the surface, we expect a change in this limit and possibly a modification of the LF–HF transfers. The study of the dependence of this limit on external parameters is left to future work, as well

as the implementation of this full energy budget in realistic eddy-resolving models.

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