Horizontal structure of nonlinear basin modes due to large scale baroclinic instability: A prototype for interannual variability

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Motivation

- In Thierry's keywords: Ocean intrinsic variability, Dynamical System Theories, Process studies idealized models, Energetics
- interannual ocean variability: atmospheric forcing but also intrinsic!
- spontaneous multidecadal variability in PE/PG models idealized configurations under constant buoyancy forcing (Huck&al. 1999 JPO, Colin de Verdière&Huck 1999 JPO, te Raa&Dijkstra 2002 JPO...), but not in 2-layer PG shallow-water models (Ferjani&al. 2013, 2014 JMR)
- large-scale baroclinic instability > potential source of low-frequency oceanic variability (Hochet&al. 2015 JPO)
- in realistic GCMs, difficult analysis of nonlinear oscillations, variance budget, eddy fluxes, linear instability (Jamet&al. 2015 CD)
- ⇒ simplest framework with large-scale baroclinic instability:
 2.5-layer shallow water model > easy analysis of local and global instability, comparison with nonlinear oscillation and eddy fluxes

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Variability of the wind-driven circulation in idealized models

interannual variability arises when Reynolds number increases $\mathrm{Re} = UL/\nu$

dynamical system approach > bifurcation structure in QG ou SW models: sensitivity to details like (free/no slip) boundary conditions, forcing symmetry... ⇒ gyre mode (Nauw&Dijkstra 2001 JMR, Simonnet&Dijkstra 2002 JPO, Dijkstra&Ghil 2005 RG...)



turbulent oscillator: positive feedback due to eddy rectification (Berloff&al. 2007 JPO)

2.5-Layer Shallow-Water Model Equations

$$\partial_t h_i + \nabla \cdot (h_i \mathbf{u_i}) = 0$$

$$\partial_t \mathbf{u_i} + \mathbf{u_i} \cdot \nabla \mathbf{u_i} + f \mathbf{k} \times \mathbf{u_i} = -\nabla \phi_i + \nu \nabla^2 \mathbf{u_i} + \delta_{i1} \frac{\tau}{\rho_i h_i}$$

$$\phi_1 = g'_{13} h_1 + g'_{23} h_2 ; \ \phi_2 = g'_{23} h_1 + g'_{23} h_2$$

$$g'_{ij} = g \frac{\rho_j - \rho_i}{\rho_i} > 0$$

 h_i layer thickness, ϕ_i hydrostatic pressure/ ρ_0 reduced gravity g'_{ij} : assumes 3rd layer has infinite depth, null circulation/pressure gradient forcing: zonal wind stress τ fct(latitude), constant in time numerical resolution in Cartesian or spherical geometry: Arakawa C-grid with energy conserving scheme (Sadourny 1975) aka "ENE" $1/3^{\circ} (1/6^{\circ})$ resolution \Rightarrow Laplacian viscosity ν 3200 (400) m² s⁻¹

Idealized configuration and forcing



Vertical stratification, basin geometry and analytical wind forcing (zonal stress and Ekman pumping), identical to Sirven&al. 2015 JMR. Control parameter: wind forcing amplitude, limited by upper layer outcropping for linearization

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A flavor of the variability: oscillations with period O(7-8 yr)



YEAR



- ▲ KE mean and standard deviation function of wind forcing amplitude > Hopf bifurcation
- ◄ Evolution of KE and APE for spherical and cartesian geometry

200

Animation of h1' anomalies: westward propagation

Influence of increasing wind forcing amplitude



(top) Standard deviation and (bottom) anomaly of h_1 (color), superposed on the mean state (contours).

Variability increases and extends further east with increasing wind forcing: link with unstable regions?

Regions of unstable waves propagation



Conditions for baroclinic instability in SW2.5



Hochet&al. 2015 JPO \rightarrow



FIG. 1. Growth time (days) of large-scale modes in a 2.5-layer quasigeostrophic model as a function of the zonal velocities in layer 1 (U_1) and layer 2 (U_2). The zonal wavelength is 1000 km, and the latitude is 30°N. The two lines are the first (dashed) and second (solid) baroclinic modes; H = 200 m.

Relation with local baroclinic instability growthrate



SW2.5 R1/3" TAUX 0.01 NU 32000 KAPPA 0 MEAN: GREEN25COMPLETUV 142200 CMEGAIMAX (YR











SW2.5 R1/3" TAUX 0.00 NU 52000 KAPPA 0 MEAN: GREEN25COMPLETUV 142200 OMEGAIMAX (YR)





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Horizontal structure of basin modes

Selection of most unstable wavelength... by eddy viscosity



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Horizontal structure of basin modes

Closed PV contours in layer 2 define instability region Rhines&Schopp 1991 JPO



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Reversed double gyre: instability at intergyre

SWN: UPPER LAYER THICKNESS STD (M) AND MEAN (CI 50M)



SW25 1/3 TAUX=-0.02: H' LAYER 1 YR = 30







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Classical double gyre with no momentum advection "PG"



at $1/3^{\circ}$, planetary geostrophic dynamics makes almost no difference with full nonlinear equations

3D linear stability analysis > technical part

• linearized version of 2.5-layer SW model (no outcropping): $X = \overline{X} + X'$, where \overline{X} is steady or time-mean state

$$\partial_t h'_i + \nabla \cdot (\overline{h_i} \mathbf{u}'_i) + \nabla \cdot (h'_i \overline{\mathbf{u}_i}) = 0$$

$$\partial_t \mathbf{u}'_i + \mathbf{u}'_i \cdot \nabla \overline{\mathbf{u}_i} + \overline{\mathbf{u}_i} \cdot \nabla \mathbf{u}'_i + f \mathbf{k} \times \mathbf{u}'_i = -\nabla \phi'_i + \nu \nabla^2 \mathbf{u}'_i - \delta_{i1} \frac{h'_i \tau}{\rho_i \overline{h_i}^2}$$

$$\phi'_1 = g'_{13} h'_1 + g'_{23} h'_2 ; \ \phi'_2 = g'_{23} h'_1 + g'_{23} h'_2$$

• computation of eigenvalues of Jacobian J through those of propagator $M(\tau)$ with ARPACK: $\partial_t X = J X = \omega X$

$$X(t=0) \to X(t=\tau) = M(\tau) X(t=0) = e^{J\tau} X(t=0),$$

- initial perturbations X(t = 0) provided by iterative Arnoldi method,

- tangent linear model integrated for 'short' time, typically $\tau{=}0.1~{\rm yr},$
- Jacobian eigenvalues ω deduced from propagator eigenvalues $e^{\omega \tau}$.

3D linear stability analysis: results



► most unstable eigenmode with growth time scale $1/\omega_r$ 76.1 yr has period 9.1 yr and wavenumber 4 similar to nonlinear oscillation

linear framework clearly separates perturbations from mean flow
 allows precise estimation of energy source and sink for modes
 sucessfully used in barotropic shallow water model (Herbette&al. 2014 GAFD) to determine regions of barotropic instability sustaining oscillation
 extended here to address baroclinic modes where APE » KE

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Energy budget > natural definition of APE

Multiplying momentum equations by $(h_i \mathbf{u_i})$:

$$\partial_t \left[h_i \frac{|\mathbf{u}_i|^2}{2} \right] + \nabla \cdot \left[h_i \mathbf{u}_i \frac{|\mathbf{u}_i|^2}{2} \right] = -h_i \mathbf{u}_i \cdot \nabla \phi_i = -\nabla (h_i \mathbf{u}_i \phi_i) - \phi_i \partial_t h_i$$

Summing last rhs term for the 2 layers leads to natural definition of APE:

$$\sum_{i} \phi_{i} \partial_{t} h_{i} = (g_{13}'h_{1} + g_{23}'h_{2}) \partial_{t} h_{1} + (g_{23}'h_{1} + g_{23}'h_{2}) \partial_{t} h_{2}$$
$$= \partial_{t} \left(g_{1}'\frac{h_{1}^{2}}{2} + g_{2}'\frac{h_{2}^{2}}{2} + g_{2}'h_{1}h_{2} \right)$$

In terms of interface deviations η_i

APE =
$$g \frac{\eta_1^2}{2} + g_{12}' \frac{\eta_2^2}{2} + g_{23}' \frac{\eta_3^2}{2}$$

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Energy budget for leading mode > exact growth rate

DIVERGENCE TERMS EKEdiv: -3.051e-17 vr⁻¹ NONLINEAR TERMS EKEV: 0.0045046 vr⁻¹ Š LATITUDE WORK OF PRESSURE FORCE EKE0: 0.10568 yr-1 WORK OF WIND × 10 EKET: 0.0018056 yr⁻¹ Š LATITUDE VISCOSITY EKEV: -0.098829 yr1 dtEEE/2EEET: 0.013163 yr-1 Š LATITUDE LONGITUDE °E -2000 -1000 2000 LONGITUDE °E

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Influence of resolution and dissipation: $1/6^{\circ}$, u=400 m² s⁻¹



The limits of crude vertical resolution



▶ in 2.5 layers, single vertical structure of unstable mode mix longwave Green type and mesoscale Charney type instability (Beckmann 1988 JPO)

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FIG. 5. Spectral distribution of maximum growth rate of unstable shear modes depending on the vertical resolution of the model. The maximum lies on the k-axis. Contour interval: 10^{-7} s⁻¹.

▶ Peculiar regime of wind-forced ocean circulation: weak forcing (1/3) and moderate dissipation \Rightarrow Reynolds number O(1–1000), not described in the litterature as far as I know?

► 'textbook case' for the analysis of local and global instability, energetics of the unstable mode...

▶ but... unstable mode wavelength set by eddy viscosity

Stronger forcing? leads to upper layer outcropping, variability regions shift to intergyre outcropping, waves become more nonlinear

Signature of these modes in oceanic variability? to be compared to response to varying wind forcing (Sirven&al. 2015 JMR)

Higher horizontal and vertical resolution? see my poster!

THANK YOU FOR YOUR ATTENTION

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